## **Precipitation Module (TC-PRISMA) User Guide**

**Thermo-Calc Version 2017b** 





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# Introduction to the Precipitation Module (TC-PRISMA)

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## **About the Precipitation Module (TC-PRISMA)**



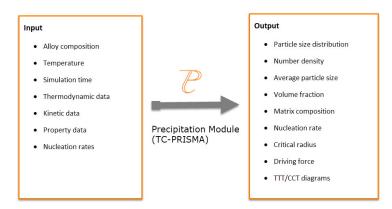
TC-PRISMA is no longer a standalone program. It is integrated into the Thermo-Calc Graphical Mode and considered an add-on module called the Precipitation Module. It is also available for all platforms (Windows, Mac and Linux). If you have older versions of the TC-PRISMA software that you want to uninstall, follow the instructions to remove this program component as described in the *Thermo-Calc Installation Guide*.

The Precipitation Module, previously referred to as TC-PRISMA, is an add-on module to the core Thermo-Calc software. The Precipitation Module itself is a general computational tool for simulating kinetics of diffusion controlled multi-particle precipitation processes in multicomponent and multiphase alloy systems.

Precipitation, formation of particles of a second phase, or second phases from a supersaturated solid solution matrix phase, is a solid state phase transformation process that has been exploited to improve the strength and toughness of various structural alloys for many years. This process is thermochemically driven and fully governed by system (bulk and interface) thermodynamics and kinetics.

Typically, a precipitation process has three distinctive stages: nucleation, growth, and coarsening. However, under certain conditions, these can also happen at the same time. With the Precipitation Module, the kinetics of concurrent nucleation, growth, and coarsening can be simulated by calculating the evolution of the probability distribution of the particle number densities, usually called particle size distribution (PSD). The simulation results can be used to understand and guide how to obtain desirable precipitates with certain PSD or to avoid undesirable precipitations during heat treatments of alloys such as aging and tempering. A summary for the input and output of the Precipitation Module is shown.

#### Input and Output of the Precipitation Module



The Precipitation Module relies on CALPHAD-based software tools and databases to provide the necessary bulk thermodynamic and kinetic data for phases in multicomponent systems. The CALPHAD approach has been developed for more than 50 years and is routinely applied to design new alloys and optimize existing materials within various metal industries, such as steels and alloys of nickel, titanium, aluminum and magnesium.

The power of this approach is due to the adopted methodology where free energy and atomic mobility of each phase in a multicomponent system can be modeled hierarchically from lower order systems, and model parameters are evaluated in a consistent way by considering both experimental data and ab-initio calculation results. The Precipitation Module is directly integrated into Thermo-Calc, a CALPHAD-based computer program for calculating phase equilibrium. Another add-on module, the Diffusion Module (DICTRA) is available for diffusion controlled phase transformation in multicomponent systems.

With Thermo-Calc and the accompanying thermodynamic and mobility databases, almost all fundamental phase equilibrium and phase transformation information can be calculated without unnecessary and inaccurate approximations. For example you can calculate:

- Driving forces for nucleation and growth
- Operating tie-lines under local equilibrium conditions
- Deviations from local equilibrium at interfaces due to interface friction
- Atomic mobilities or diffusivities in the matrix phase

In addition to bulk thermodynamic and kinetic data, a few other physical properties, such as interfacial energy and volume, are needed in precipitation models. These additional physical parameters can be obtained by experiments or other estimation models or first principles calculations. Volume data for steels and nickel-based alloys have already been assessed and included in TCFE, TCNI, and TCAL databases. The Precipitation Module has an estimation model available for interfacial energy.



This guide is a supplement to the full Thermo-Calc documentation set. It is recommended that you use the Online Help: From the main menu in Thermo-Calc choose  $Help \rightarrow Online \ help$ .



See "Help Resources" on the next page to learn how to access this information if you have not already done so.

## **Help Resources**

#### 'Online' Help: No Internet Required

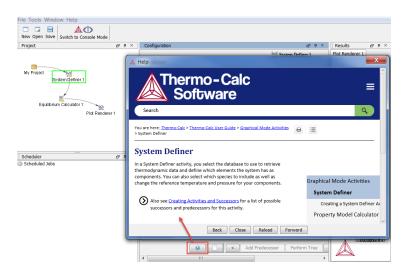
To access online help in a browser, open Thermo-Calc and select  $Help \rightarrow Online Help$ . The content opens in a browser but uses local content so you don't need an Internet connection except for links to external websites.

#### **Context Help (Graphical Mode Only)**



When you are in Graphical Mode, you can access feature help (also called *topic-sensitive* or *context help*) for the activity nodes in the tree.

- 1. In the Project window, click a node. For example, **System Definer**.
- 2. In the lower left corner of the **Configuration** window, click the help button .
- 3. The **Help** window opens to the relevant topic.



The window that opens has the same content as the help you access in the browser. There are these extra navigation buttons in this view.

• Use the buttons on the bottom of the window, Back Close Reload Forward, to navigate Back and Forward (these are only active once you have started using the help to load pages and create a history), to Close the window, and Reload the original content.

### **Console Mode Help**

Console Mode is for Thermo-Calc and the Diffusion Module (DICTRA).

In Console Mode at the command line prompt, you can access help in these ways:

- For a list of all the available commands in the current module, at the prompt type a question mark (?) and press <Enter>.
- For a description of a specific command, type Help followed by the name of the command. You can only get online help about a command related to the current module you are in.
- For general system information type Information. Specify the subject or type ? and the available subjects are listed. This subject list is specific to the current module.

## **Available Options**

The Precipitation Module, previously referred to as TC-PRISMA, is an add-on module to the core Thermo-Calc software. A separate license is required to perform calculations for more than three elements. Without it you are able to use the module in *Demo Mode*.

"Demonstration (Demo) Mode" on the next page

#### **Precipitation Simulation Template**

A **Precipitation Simulation** template is available to all Thermo-Calc users.

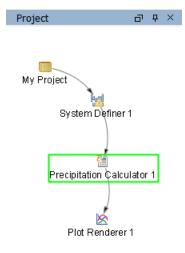


If you are accessing the Precipitation Module in Demo Mode, see "Demonstration (Demo) Mode" on the next page for what is available to you.

#### **USING THE TEMPLATE**

When you first open Thermo-Calc in Graphical Mode, on the **My Project** window, click the **Precipitation Simulation** button to add a *System Definer, Precipitation Calculator* and *Plot Renderer* to the **Project** tree.

Click the **Precipitation Simulation** button in the templates section to add these nodes to the Project.



"Creating a "Project from a Template" in the Thermo-Calc User Guide or search the Online Help.

#### **Precipitation Calculator**

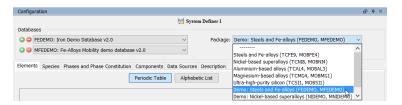
A Precipitation Calculator allows you to set the conditions for, and perform, a precipitation calculation. Details about this activity are described in Precipitation Calculator.

#### **Demo Database Packages for the Examples Collection**

Both a thermodynamic and mobility database are required to run the Precipitation Module simulation. If you have a Precipitation Module (TC-PRISMA) license you can run all the examples as the demonstration database packages are included with your installation. Select the database packages from the **System Definer**—**Configuration** window to run a simulation.



If you are using the Precipitation Module in Demo Mode, see "Demonstration (Demo) Mode" below to learn more.



### **Demonstration (Demo) Mode**

The Precipitation Module, and some examples (see "Precipitation Module Examples" on page 15), are available to all Thermo-Calc users but only for simulations with three elements. If you do not have a license for the Precipitation Module then you are in *Demonstration Mode* when using the Precipitation Calculator or Precipitation Simulation template.

"Network License Restrictions" on page 13

#### PRECIPITATION SIMULATION TEMPLATE

When you are in DEMO mode, in the **Templates** area this is indicated by the text under the logo.



#### PRECIPITATION CALCULATOR

If you are experimenting with the Precipitation Calculator in Demo Mode, you may have access to a variety of databases based on your license. However, you can only define three elements for a demo simulation.

If you define more than three elements on the System Definer, when you go to the Precipitation Calculator, the **Perform** button is unavailable and the tooltip explains why. In this case one of the chosen elements needs to be removed and then the Perform button is made available.





Even if you have more than three elements, the Plot Renderer or Table Renderer **Perform** button is still available. However, if you click the button and try to run the simulation the Event Log displays an error message.

#### **Network License Restrictions**

The Precipitation Module (TC-PRISMA) requires a separate license. If you are using a network client installation of Thermo-Calc, then you may not be able to use it even if you have access to a license server with a valid network license file. The reason for this is because other clients who are part of your network installation may have checked out all instances of the network license allowed to run simultaneously.



For users with a network license, you must exit Thermo-Calc to release the license for other users. The license is checked out as soon as you add a Precipitation Calculator and remains unavailable to other users until you exit the program.

- With a network license, and if as per above you temporarily do not have access to a license, you are automatically put into *Demo Mode*. Then the Precipitation Calculator is available with three elements.
- Search the online help or see the *Thermo-Calc Installation Guide* for more about network licenses.

## Selecting the Disordered Phase as a Matrix Phase

The following information is about using disordered Fcc as a matrix phase with the following thermodynamic and mobility database packages:

- TCAL and MOBAL (Al-based alloys)
- TCCU and MOBCU (Cu-based alloys)
- TCNI and MOBNI (Ni-based alloys)

In the TCNI/MOBNI, TCAL/MOBAL, and TCCU/MOBCU packages, the well-known order/disorder two-sublattice model is used to describe the Gibbs energy of both FCC\_A1 and FCC\_L12. With this treatment, FCC\_L12 is becoming FCC\_A1 if the site fractions of each element on both sublattices are identical, which means that FCC\_A1 is only a special case of FCC\_L12. Therefore, FCC\_A1 is not shown in the phase list on the *Phases and Phase Constitution* tab on the System Definer activity and in subsequent equilibrium calculation results. Instead it is shown only as FCC\_L12. The real ordered FCC\_L12 is shown as FCC\_L12#2.

In precipitation simulations, the matrix phase is quite often the disordered FCC phase. You can directly select FCC\_L12 as the matrix phase and run a simulation. However, the speed is not optimal due to the sophisticated model used for both Gibbs energy and atomic mobilities. A better and more convenient way is to deselect FCC\_L12 and FCC\_L12#2 from the phase list on the *Phases and Phase Constitution* tab on the **System Definer** if the ordered phase is irrelevant in the alloy under investigation, such as in most Al and Cu alloys. Once these are unchecked (i.e. not selected), the FCC\_A1 phase is available and can later be selected as the matrix phase.

For Ni-based superalloys using the TCNI/MOBNI package, the ordered FCC\_L12#2 (gamma prime) has to be included as the precipitate phase in most of calculations. In this case, you can select DIS\_FCC\_A1 from the phase list on the *Phases and Phase Constitution* tab and then select it as the matrix phase in the **Precipitation Calculator**.

## **Precipitation Module Examples**



Examples that use three elements are available to all users. The other examples require a Precipitation Module (TC-PRISMA) license to calculate and plot results.



All examples use demonstration database packages included with your installation. You can open the examples from the main menu: File → Open Examples or Help → Open Examples.



Unless specified in tables for each example, all the numerical parameters are assumed default values.

## P\_01: Isothermal Precipitation of Al<sub>3</sub>Sc

This example simulates the kinetics of precipitation of  $Al_3Sc$  from an FCC\_A1 solution phase. The simulation results can be compared with experimental data collected from Marquis and Seidman<sup>1.</sup> and Novotny and Ardell (2001)<sup>2.</sup>.



This example is available as a video tutorial on our website and YouTube channel.

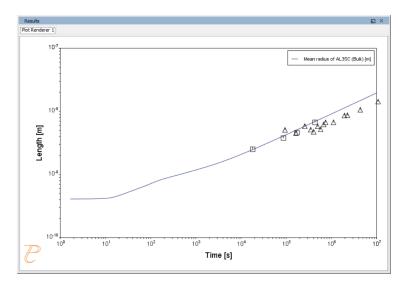
System (System Definer)		
Database package	Database package Demo: Aluminum-based alloys (ALDEMO, MALDEMO)	
Elements	Al, Sc	
Conditions (Precipitation Calculator)		
Composition	Al-0.18Sc Mole percent	
Matrix phase	FCC_A1	

<sup>1.</sup> Marquis, E.A, and D.N Seidman. 2001. "Nanoscale Structural Evolution of Al3Sc Precipitates in Al(Sc) Alloys." Acta Materialia 49 (11): 1909–19.

<sup>&</sup>lt;sup>2.</sup> Novotny, Gabriel M., and Alan J. Ardell. 2001. "Precipitation of Al3Sc in Binary Al–Sc Alloys." Materials Science & Engineering, A: Structural Materials: Properties, Microstructure and Processing 318 (1–2): 144–54.

Precipitate phase	AL3SC	
Precipitate Phase Data Parameters (Precipitation Calculator)		
Nucleation sites	Bulk	
Interfacial energy	Calculated	
Calculation Type (Precipitation Calculator)		
Calculation type	Isothermal	
Temperature	350° C	
Simulation time	1.0E7 seconds	
Other		
File name and location	Folder: Precipitation Module - TC-PRISMA  File name: P_01_Precipitation_Al-Sc_AL3SC.tcu	

#### Mean radius



### P\_02: Stable and Metastable Carbides - Isothermal



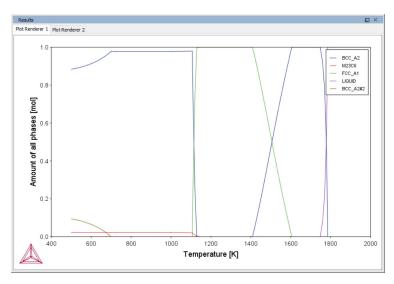
This example simulates the kinetics of precipitation of both stable and metastable carbides from ferrite phase. It demonstrates that metastable carbides (cementite, M7C3) may first emerge and then disappear and the stable phase (M23C6) prevails.

This example uses the Equilibrium Calculator and a property diagram to determine how the phases change with temperature. We are interested in the carbide precipitation at 1053K where only the carbide M23C6 is stable according to the equilibrium calculation. The Precipitation Calculator is used to do an isothermal calculation of the three phases (cementite, M23C6 and M7C3) where cementite and M7C3 are metastable phases.

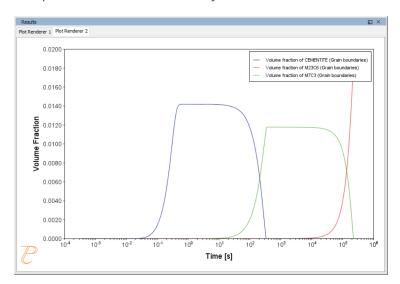
System (System Definer)		
	Domas Steele and Fe alleve (FFDFMO MFFDFMO)	
Database package	Demo: Steels and Fe-alloys (FEDEMO,MFEDEMO)	
Elements	Fe, C, Cr	
Conditions (Precipitation Calculator)		
Composition	Fe-0.1C-12Cr Mass percent	
Matrix phase	BCC_A2	
Precipitate phases	Cementite, M23C6 and M7C3	
Matrix Phase Data Parameters (Precipitation Calculator)		
Grain size (click <b>Show details</b> to display this setting)	1.0E-4 m	
Precipitate Phase Data Parameters (Precipitation Calculator)		
Nucleation sites	Grain boundaries	
Interfacial energy	Cementite 0.167 J/m <sup>2</sup> , M23C6 0.252 J/m <sup>2</sup> , M7C3 0.282 J/m <sup>2</sup>	
Calculation Type (Precipitation Calculator)		
Calculation type	Isothermal	
Temperature	1053 K	

Simulation time	400 000 seconds	
Other		
File name and location	Folder: Precipitation Module - TC-PRISMA  File name: P_02_Precipitation_Fe-C-Cr_Cementite-M7C3- M23C6.tcu	

Equilibrium Calculator: How the phases change with temperature



#### Precipitation Calculator: Volume fractions



#### P\_03: Stable and Metastable Carbides - TTT Diagram



In this example, the kinetics of precipitation of both stable and metastable carbides is calculated from the ferrite phase. It demonstrates that metastable carbides may first emerge and then disappear and the stable phase prevails.

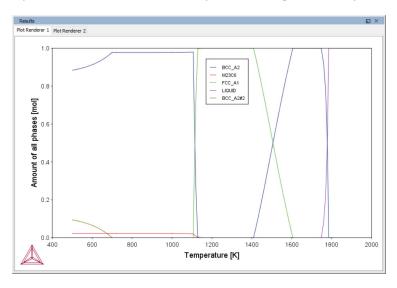
This example uses the Equilibrium Calculator and a property diagram to determine how the phases change with temperature. Using this result, the Precipitation Calculator is used to do a TTT (Time-Temperature-Transformation) diagram calculation of the three phases (cementite, M23C6 and M7C3) at the grain boundaries.

For a TTT diagram calculation, select **TTT diagram** in **Calculation Type**, then enter **Min**, **Max**, and **Step** of **Tempeature**, as well as **Max annealing time**. In **Stop criterion**, choose **Volume fraction of phase** and enter the value.

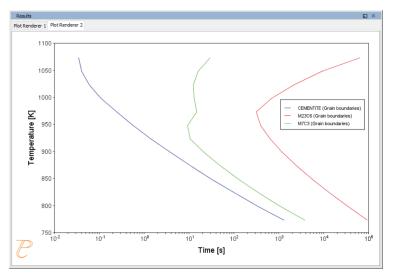
System (System Definer)		
Database package	Demo: Steels and Fe-alloys (FEDEMO and MFEDEMO)	
Elements	Fe, C, Cr	
Conditions (Precipitation Calculator)		
Composition	Fe-0.1C-12Cr Mass percent	
Matrix phase	BCC_A2	
Precipitate phases	Cementite, M23C6 and M7C3	
Matrix Phase Data Parameters (Precipitation Calculator)		
Grain size (click <b>Show details</b> to display this setting)	1.0E-4 m	
Precipitate Phase Data Parameters		
Nucleation sites	Grain boundaries	
Interfacial energy	Cementite 0.167 J/m <sup>2</sup> , M23C6 0.252 J/m <sup>2</sup> , M7C3 0.282 J/m <sup>2</sup>	
Calculation Type (Precipitation Calculator)		
Calculation type	TTT diagram	

Temperature	500° to 800° C with 25° C steps	
Max. annealing time	1.0E8 seconds	
Stop criteria	Volume fraction of phase is set to 0.0001	
Other		
File name and location	Folder: Precipitation Module - TC-PRISMA  File name: P_03_Precipitation_Fe-C-Cr_TTT_Cementite-M7C3- M23C6.tcu	

Equilibrium Calculator: How the phases change with temperature



### Precipitation Calculator: TTT diagram



## P\_04: Precipitation of Iron Carbon Cementite

Simulate the kinetics of precipitation of carbides from a BCC Fe solution phase. This example is based on Wert (1949)<sup>1</sup>.

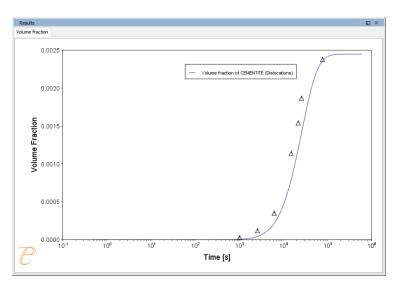
This example uses two Experimental File Reader activities with the Precipitation Calculator and an isothermal calculation, to plot the volume fraction of the cementite phase.

System (System Definer)		
Database package	Demo: Steels and Fe-alloys (FEDEMO and MFEDEMO)	
Elements	Fe, C	
Conditions (Precipitation Calculator)		
Composition	Fe-0.016C mass percent	
Matrix phase	BCC_A2	
Precipitate phase	Cementite	
Matrix Phase Data Parameters (Precipitation Calculator)		
Grain aspect ratio (click <b>Show details</b> to display this	1.0	

<sup>&</sup>lt;sup>1.</sup> Wert, Charles A. 1949. "Precipitation from Solid Solutions of C and N in  $\alpha$ -Iron." Journal of Applied Physics 20 (10). AIP Publishing: 943.

setting)		
Dislocation density (click <b>Show details</b> to display this setting)	1.5e11m <sup>-3</sup>	
Precipitate Phase Parameters (Precipitation Calculator)		
Nucleation sites	Dislocations	
Interfacial energy	0.24 J/m <sup>2</sup>	
Calculation Type (Precipitation Calculator)		
Calculation type	Isothermal	
Temperature	102° C	
Simulation time	600 000 seconds	
Options (Precipitation Calculator)		
Growth rate model	Advanced	
Other		
File name and location	Folder: Precipitation Module - TC-PRISMA  File name: P_04_Precipitation_Fe-C_Cemetite.tcu	

#### Volume Fraction



### P\_05: Precipitation of $\gamma'$ in Ni Superalloys - Isothermal



This example simulates the kinetics of precipitation of  $\gamma'$  phase from  $\gamma$  phase. The simulation results can be compared with experimental data collected from Sudbrack et al.  $(2008)^1$ .

This example uses three Experimental File Reader activities with the Precipitation Calculator. It does an isothermal calculation to plot the volume fraction, mean radius and number density of the cementite phase.



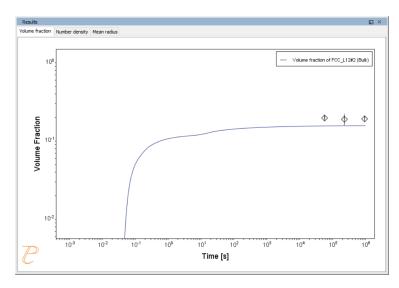
DIS\_FCC\_A1 needs to be selected on the System Definer. Search the online help for **Selecting the Disordered Phase as a Matrix Phase** in the *Precipitation Module* (*TC-PRISMA*) *User Guide* for details.

System (System Definer)	
Database package	Demo: Nickel-based Super Alloys (NIDEMO and MNIDEMO)
Elements	Ni, Al Cr
Conditions (Precipitation Calculator)	
Composition	Ni-9.8Al-8.3Cr Mole percent
Matrix phase	DIS-FCC_A1 (see note above about how to select this phase)
Precipitate phase	FCC_L12#2
Precipitate Phase Data Parameters (Precipitation Calculator)	
Nucleation sites	Bulk
Interfacial energy	0.012 J/m <sup>2</sup>
Calculation Type (Precipitation Calculator)	
Calculation type	Isothermal
Temperature	800° C

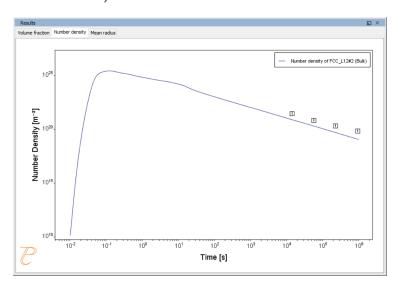
<sup>&</sup>lt;sup>1.</sup> Sudbrack, Chantal K., Tiffany D. Ziebell, Ronald D. Noebe, and David N. Seidman. 2008. "Effects of a Tungsten Addition on the Morphological Evolution, Spatial Correlations and Temporal Evolution of a Model Ni–Al–Cr Superalloy." Acta Materialia 56 (3): 448–63.

Simulation time	1 000 000 seconds
Other	
File name and location	Folder: Precipitation Module - TC-PRISMA  File name: P_05_Precipitation_Ni-Al-Cr_Isothermal_Gamma-Gamma_ prime.tcu

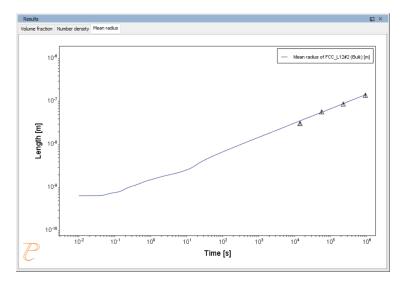
#### **Volume Fraction**



#### Number Density



#### Mean Radius



## P\_06: Precipitation of $\gamma'$ in Ni Superalloys - Non-isothermal

P

This example simulates the kinetics of precipitation of  $\gamma'$  phase from  $\gamma$  phase in Ni-8Al-8Cr and Ni-10Al-10Cr at.% alloys during continuous cooling. The simulation results can be compared with experimental results from Rojhirunsakool et al.  $(2013)^1$ .

In these examples a linear, continuously cooling condition is set up. This is done on the Precipitation Calculator settings window in the **Calculation Type** section:

- 1. Select a Non-Isothermal.
- 2. Select a **Temperature Unit** and **Time Unit** from the lists.
- 3. Click **Thermal Profile**, where **Time-Temperature** profile points are entered to describe the stepwise heat treatment schedule with linear segments. Up to 999 rows of data points can be entered.
  - In these examples two rows of data points are needed for starting time 0s and temperature 1150° C, as well as end time 3300s and temperature 380° C.
- 4. Click the **Plot Renderer** node to select the **Separate multimodal PSD** check box. This plots the mean radius and size distributions of the two compositions.

<sup>&</sup>lt;sup>1.</sup> Rojhirunsakool, Tanaporn, S. Meher, J. Y. Hwang, S. Nag, J. Tiley, and Rajarshi Banerjee. 2013. "Influence of Composition on Monomodal versus Multimodal Γ' Precipitation in Ni–Al–Cr Alloys." Journal of Materials Science 48 (2): 825–31.

Plotting the size distribution from the final simulation time of 3300 seconds, you can see there are several peaks, although these are not completely separated. As above, select the **Separate multimodal PSD** check box to separate the peaks.

 Adjust the Valley depth ratio setting to 0.05 to separate into two peaks as shown in the Ni-10Al-10Cr plot example. You can experiment with this setting to see how the size distribution evolves with time, for example, try entering several values as plot times 400 600 3300.

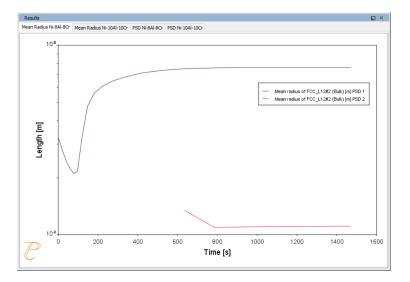


DIS\_FCC\_A1 needs to be selected on the System Definer. Search the online help for **Selecting the Disordered Phase as a Matrix Phase** in the *Precipitation Module* (TC-PRISMA) User Guide for details.

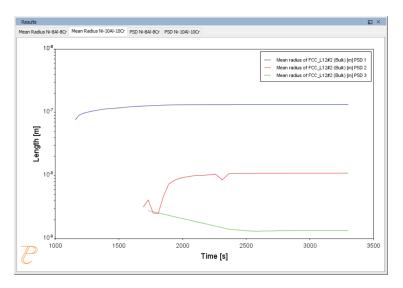
System (System Definer)		
Database package	Demo: Nickel-based Super Alloys (NIDEMO and MNIDEMO)	
Elements	Ni, Al, Cr	
Conditions (Precipitation Calculator)		
Composition (Ni-8Al-8Cr)	Ni-8Al-8Cr Mole percent	
Composition (Ni-10Al-10Cr)	Ni-10Al-10Cr Mole percent	
Matrix phase	DIS_FCC_A1 (see note above about how to select this phase)	
Precipitate phase	FCC_L12#2	
Matrix Phase Data Parameters (Precipitation Calculator)		
Mobility enhancement prefactor (click Show details to display this setting)	5.0	
Precipitate Phase Data Parameters (P	recipitation Calculator)	
Nucleation sites	Bulk	
Interfacial energy	0.023 J/m <sup>2</sup>	
Calculation Type (Precipitation Calculator)		
Calculation type	Non-isothermal	

Temperature unit	Celsius	
Time unit	Seconds	
Simulation time (Ni-8Al-8Cr)	1470 s	
Simulation time (Ni-10Al-10Cr)	3300 s	
Multimodal PSD (Plot Renderer)		
Separate multimodal PSD for 8Al-8Cr	The <b>Valley depth ratio</b> is set to 0.05 for both plots. The number of <b>Points</b> is increased to 200 for an average radius plot.	
Separate multimodal PSD for 10Al-10Cr	The <b>Valley depth ratio</b> is set to 0.18 for both plots.	
Other		
File name and location	Folder: Precipitation Module - TC-PRISMA  File name: P_06_Precipitation_Ni-Al-Cr_Non-isothermal_Gamma- Gamma_prime.tcu	

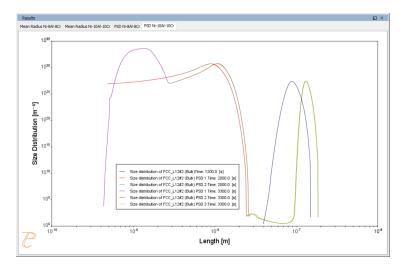
Mean Radius Ni-8Al-8Cr



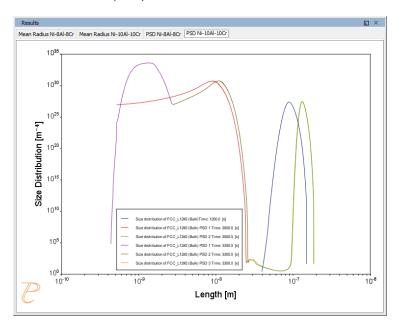
#### Mean Radius Ni-10Al-10CR



#### Size Distribution (PSD) Ni-8Al-8Cr



#### Size Distribution (PSD) Ni-10Al-10Cr



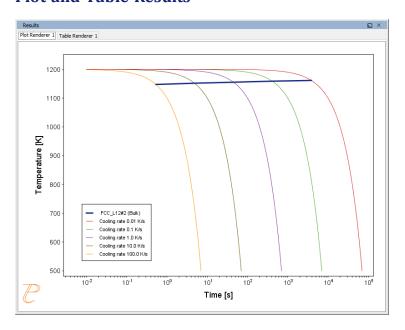
## P\_07: Cooling Rate (CCT) Diagram of Ni-Al-Cr $\gamma$ - $\gamma$ '

In this example, a CCT diagram for a Ni-10Al-10Cr  $\gamma$ - $\gamma'$  alloy is calculated and plotted with superimposition of the cooling rate values.

System (System Definer)	
Database package	Demo: Nickel-based Super Alloys (NIDEMO and MNIDEMO)
Elements	Ni, Al, Cr
Conditions (Precipitation Calculator)	
Composition	Ni-10Al-10Cr Mole percent
Matrix phase	DIS_FCC_A1
Precipitate phase	FCC_L12#2
Precipitate Phase Data Parameters (Precipitation Calculator)	
Nucleation sites	Bulk

Interfacial energy	0.023 J/m <sup>2</sup>	
Calculation Type (Precipitation Calculator)		
Calculation type	CCT Diagram	
Temperature Min to Max	500 to 1200 Kelvin	
Cooling rate(s)	.01 .1 1 10 100 K/s	
Stop criteria	Volume fraction of phase 1.0E-4	
Other		
File name and location	Folder: Precipitation Module - TC-PRISMA  File name: P_07_Precipitation_Ni-Al-Cr_CCT_Gamma-Gamma_prime.tcu	

#### **Plot and Table Results**



Results	
Plot Renderer 1 Tab	le Renderer 1
Temperature [K]	FCC_L12#2 (Bulk)
1147.52034	0.53291
1151.74192	4.83553
1155.24636	44.76534
1158.85368	411.49173
1161.48750	3851.47779

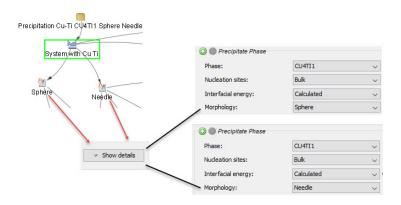
## P\_08: Precipitation of Cu-Ti CU4TI1 with Assumptions of Sphere and Needle Morphologies

In this isothermal calculation example, the precipitation of Cu4Ti phase in a Cu-Ti binary alloy is calculated. To make a comparison, two separate simulations are performed, one assuming spherical morphology without elastic strain energy, and the other assuming needle morphology whose shape, determined by competition between interfacial energy and elastic strain energy, is changed during the simulation. The transformation strain is obtained from Borchers<sup>1</sup>. The results are compared with experiment results from Kampmann et al<sup>2</sup>.



To ensure that the settings are done on the correct Precipitation Calculators, the **Sphere** and **Needle** nodes are renamed from **Precipitation Calculator** to match their morphology. The morphology is set in the **Precipitate Phase** section when you click **Show details**.

#### "Precipitation Morphology" on page 89



System (System Definer)	
Database package	Demo: Copper-based alloys (CUDEMO and MCUDEMO)
Elements	Cu, Ti
Sphere and Needle Conditions (Precipitation Calculator)	
Composition	Cu-1.9Ti Mole percent

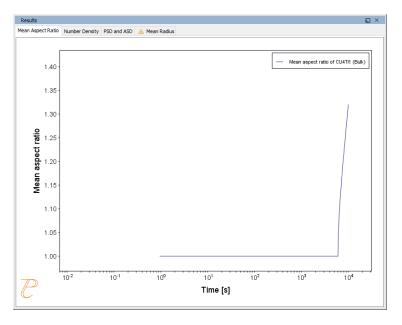
<sup>1.</sup> Borchers C. 1999. "Catastrophic Nucleation during Decomposition of Cu-0.9at.%Ti." Phil. Mag. A 79(3):537-547

<sup>&</sup>lt;sup>2.</sup> Kampmann R., Eckerlebe H., and Wagner R. 1987. "Precipitation Kinetics in Metastab le Solid Solutions - Theoretical Considerations and Application to Cu-Ti Alloys." Mat. Res. Soc. Symp. Proc. 57: 525-542.

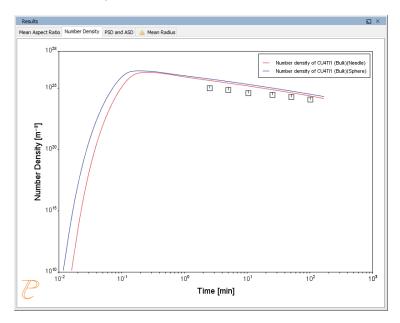
Matrix phase	FCC_L12
Precipitate phase	CU4TI1
Matrix Phase Data Parameter	s (Precipitation Calculator)
Mobility enhancement prefactor (click <b>Show details</b> to display this setting)	100
Precipitate Phase Data Param	eters (Precipitation Calculator)
Nucleation sites	Bulk
Interfacial energy	The default
Morphology (click <b>Show details</b> to display this setting)	For the <b>Sphere</b> node (renamed from Precipitation Calculator), keep the default.  For the <b>Needle</b> node (renamed from Precipitation Calculator), <b>Needle</b> is selected.
Transformation strain (click <b>Show details</b> to display this setting)	For the <b>Sphere</b> node (renamed from Precipitation Calculator), keep the default.  For the <b>Needle</b> node (renamed from Precipitation Calculator), <b>User defined</b> is selected. In this example, the following settings are defined:  • £11 and £22 are set to <b>0.022</b> • £33 is set to <b>0.003</b>
Calculation Type (Precipitatio	n Calculator)
Calculation type	Isothermal
Temperature	350° C

Simulation time	10,000 seconds	
Datasets (Experimental File Reader)		
Borchers Mean radius vs Time and Borchers Number density vs Time	Data sets included with this example and imported to two Experimental File Readers. These data sets are used for the Mean Radius and Number Density plots, respectively.	
Other		
File name and location	Folder: Precipitation Module - TC-PRISMA  File name: P_08_Precipitation_Cu-Ti_CU4TI1_Sphere_Needle.tcu	

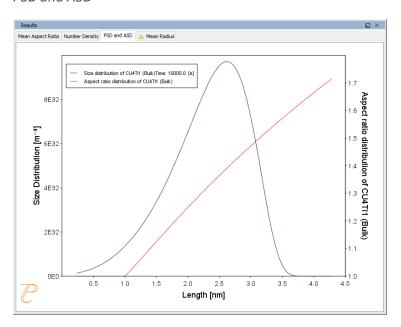
#### Mean Aspect Ratio



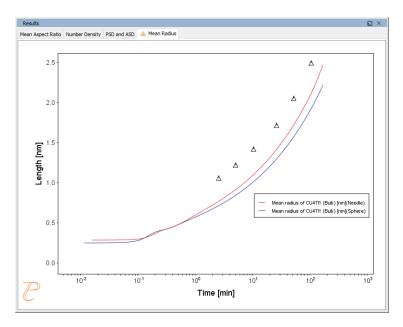
#### Number Density



#### PSD and ASD



#### Mean Radius



## P\_09: Precipitation of Al-Sc AL3SC with Assumption of Sphere and Cuboid Morphologies

In this isothermal calculation example, the precipitation of Al<sub>3</sub>Sc phase from FCC\_A1 matrix phase in a Al-Sc binary alloy is simulated. To make a comparison, two separate calculations are performed, one assuming spherical morphology without elastic strain energy, and the other assuming cuboid morphology whose shape is determined by competition between interfacial energy and elastic strain energy. The simulation results are compared with experimental data collected from Marquis and Seidman<sup>1.</sup> and Novotny and Ardell (2001)<sup>2.</sup> In addition, mean cubic factor and cubic factor distribution are also plotted for cuboid shape to illustrate the spherical-cuboidal transition during precipitation.



To ensure that the settings are done on the correct Precipitation Calculators, the **Sphere** and **Cuboid** nodes are renamed from **Precipitation Calculator** to match their morphology. The morphology is set in the **Precipitate Phase** section when you click **Show details**. See P 08 for an example of this.

"Precipitation Morphology" on page 89

<sup>1.</sup> Marquis, E.A, and D.N Seidman. 2001. "Nanoscale Structural Evolution of Al3Sc Precipitates in Al(Sc) Alloys." Acta Materialia 49 (11): 1909–19.

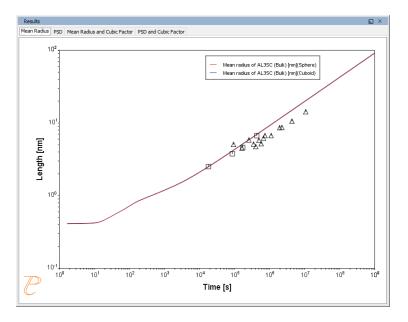
<sup>&</sup>lt;sup>2.</sup> Novotny, Gabriel M., and Alan J. Ardell. 2001. "Precipitation of Al3Sc in Binary Al–Sc Alloys." Materials Science & Engineering, A: Structural Materials: Properties, Microstructure and Processing 318 (1–2): 144–54.

Database package	Demo: Aluminum-based alloys (ALDEMO, MALDEMO)
Elements	Al, Sc
Sphere and Cuboid C	onditions (Precipitation Calculator)
Composition	Al-0.18Sc Mole percent
Matrix phase	FCC_A1
Precipitate phase	AL3SC
Matrix Phase Data Pa	arameters (Precipitation Calculator)
	For the <b>Sphere</b> node (renamed from Precipitation Calculator), the default, <b>Disregar</b> is kept.
Elastic properties (click <b>Show details</b> to	For the <b>Cuboid</b> node (renamed from Precipitation Calculator), choose <b>Cubic</b> . Then enter the elastic constants accordingly. Default elastic constants are given based on the major element of the alloy system. In this example that is
display this setting)	• c11 is <b>108.2</b> GPa
	• c12 is <b>61.3</b> GPa
	• c44 is <b>28.5</b> GPa
Precipitate Phase Dat	ta Parameters (Precipitation Calculator)
Nucleation sites	Bulk
Interfacial energy	The default
Morphology (click	For the <b>Sphere</b> node (renamed from Precipitation Calculator), keep the default.
<b>Show details</b> to display this setting)	For the <b>Cuboid</b> node (renamed from Precipitation Calculator), <b>Cuboid</b> is selected.
Transformation strain (click <b>Show details</b> to display this setting)	For the <b>Sphere</b> node (renamed from Precipitation Calculator), keep the default.
	For the <b>Cuboid</b> node (renamed from Precipitation Calculator), <b>Calculate from mola volume</b> is selected to obtain a purely dilatational strain.
Calculation Type (Pre	ecipitation Calculator)
Calculation type	Isothermal

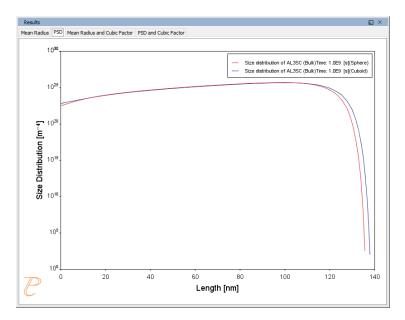
Temperature	350° C			
Simulation time	1.0E9 seconds			
Datasets (Experimental File Reader)				
Dataset 1 and Dataset 2	Data sets included with this example and imported to one Experimental File Reader. It is used for the Mean Radius plot.			
Other				
File name and location	Folder: Precipitation Module - TC-PRISMA  File name: P_09_Precipitation_Al-Sc_AL3SC_Sphere_Cuboid.tcu			

# **Plot Results**

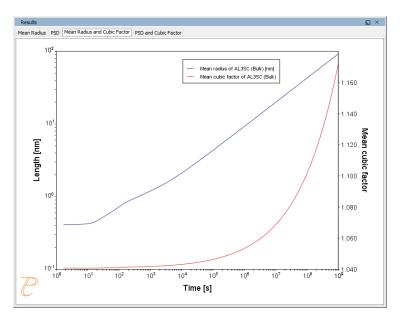
### Mean Radius



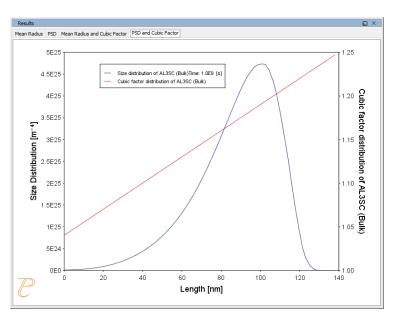
## PSD



### Mean Radius and Cubic Factor



## PSD and Cubic Factor



# **Using the Precipitation Calculator**

## In this section:

Precipitation Calculator	41
Configuration Settings	42
Plot Renderer	51
Plot Settings	51
Non-Isothermal Simulations	54

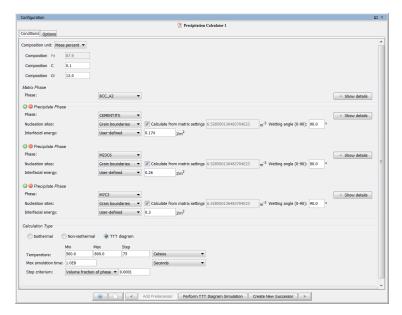
# **Precipitation Calculator**

A Precipitation Calculator allows you to set the conditions for, and perform, a precipitation calculation. The Configuration window for a Precipitation Calculator has these tabs where you define the settings:

- **Conditions**: Set the conditions for your calculation that define the Matrix and Precipitate phases. Choose the Calculation Type.
- **Options**: Modify Numerical Parameters that determine how the conditions are calculated. The Growth rate model can be set to Simplified or Advanced mode.

# **Configuration Settings**

Below are details about the settings available from the **Conditions** and **Options** tabs.

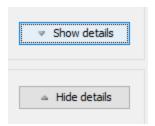




You can edit the defaults for some of these settings by going to **Tools**  $\rightarrow$  **Options**  $\rightarrow$  **Graphical Mode**  $\rightarrow$  **Precipitation**.

## Show or Hide Details

Click **Show details** (found to the right of the section) to view some of the additional settings listed below.



# Composition unit

Choose from Mass percent, Mole percent, Mass fraction or Mole fraction.

#### Matrix Phase

#### Phase



Only phases with kinetic data can be selected as the matrix phase. If the list is empty, go to the System Definer to make sure that both thermodynamic and kinetic databases are selected and defined. See "System Definer" in the *Thermo-Calc User Guide* or search the Online Help.

Choose a phase from the list. The list is based on the settings for the System Definer. When setting up your system, choose a matrix phase with kinetic data available in the database.

#### **Elastic properties**

Choose the default **Disregard** to ignore the elastic properties. Otherwise choose **Isotropic** or **Cubic** and enter the elastic constants accordingly. Default elastic constants are given based on the major element of your alloy system.

#### Molar volume

Choose the default to take the value from the **Database** or **User defined** to enter another value. **Database** is available in the list when the molar volume for the phase is defined in the thermodynamic database.

#### **Grain size**

Enter a numerical value and choose a unit from the list. The default is 1.0E-4 m.

#### **Grain aspect ratio**

Enter a numerical value. The default is 1.0.

#### **Dislocation density**

Enter a numerical value. The default is 5.0E12 m<sup>-3</sup>.

#### Mobility enhancement prefactor

A parameter that multiplies to the mobility data from database.

#### Mobility enhancement activation energy

A value that adds to the activation energy of mobility data from database.

#### **Precipitate Phase**

#### **Phase**



The phases available to choose have both thermodynamic and kinetic data. If the list is empty, go to the System Definer to make sure that both types of databases are selected and defined. See "System Definer" in the *Thermo-Calc User Guide* or search the Online Help.

Choose a phase from the list. The list is based on the settings for the System Definer.

#### **Nucleation sites**



Choose one of the following from the list.

- Bulk, Grain boundaries, Grain edges, Grain corners, or Dislocations.
- For Grain boundaries, Grain edges and Grain corners, enter the Wetting angle in addition to the matrix settings.

Click to select the **Calculate from matrix settings** check box if you want to calculate the number density of sites from the matrix grain size or dislocation density.

To enter a specific value for the number of **Nucleation sites**, deselect the check box.

#### **Interfacial energy**



Choose **Calculated** to use the estimated value and then enter a different **prefactor** value if you want to adjust the estimated value. You can also choose **User defined** to enter a value in J/m<sup>2</sup>.

### Morphology

"Precipitation Morphology" on page 89or search the online help.

Choose the particle shape from the list: **Sphere** (default), **Cuboid**, **Plate** and **Needle**.

**Cuboid** is available only when **Cubic** is chosen in **Elastic properties** of Matrix Phase.

When Plate or Needle is chosen, a check box is provided to Calculate aspect ratio from elastic energy, which is available when Isotropic or Cubic is chosen in Elastic properties of Matrix Phase. Alternatively a value can be entered into the **Aspect ratio** field to provide a constant aspect ratio.

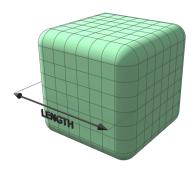


Figure 1: Cuboids have six faces, which form a convex polyhedron.

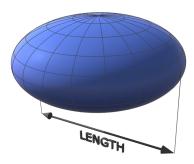


Figure 2: Oblate spheroids have rotational symmetry around an axis from pole to pole.

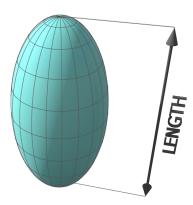


Figure 3: A prolate spheroid is a surface of revolution obtained by rotating an ellipse about its major axis.

#### **Transformation strain**

"Precipitation Morphology" on page 89

Choose the default, **Disregard**, to ignore the transformation strain. Otherwise choose **Calculated from molar volume** to obtain a purely dilatational strain, or **User defined** to enter the values.

These options are available only when **Isotropic** or **Cubic** is chosen in **Elastic properties** in Matrix Phase. In addition, if **Cuboid** is chosen as **Morphology**, only **Calculated from molar volume** is available.

#### Molar volume

Choose the default to take the value from the **Database** or **User defined** to enter another value. **Database** is available in the list when the molar volume for the phase is defined in the thermodynamic database.

#### Phase boundary mobility

A parameter that accounts for interface-controlled growth. Only effective if a very small, positive value is used. Use with caution due to a tentative treatment.

#### Phase energy addition

An energy value that adds to the Gibbs free energy from database.

#### **Approximate driving force**

> "Nucleation Theory" on page 61

Select the check box to include this if simulations with several compositions sets of the same phase create problems.

### Calculation Type

#### **Isothermal**

Use an **Isothermal** calculation type to do a precipitation simulation at constant temperature. Enter a **Temperature** and **Simulation time**.



See examples P\_01, P\_02, P\_04, P\_05, P\_08 and P\_09 in the *Graphical Mode Examples Guide*.

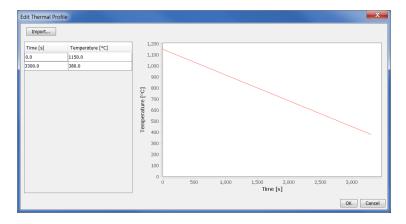
#### Non-isothermal



"Nucleation During a Non-isothermal Process" on page 74

For the **Non-isothermal** calculation type, select a **Temperature unit** and **Time unit** from the lists. Enter a value for the **Simulation time**.

Click **Thermal Profile**. Here the **Temperature** and **Time** coordinates of thermal profile points are entered. A minimum of two points is required. You can also click **Import** to add your own thermal profile from an Excel spreadsheet.



4:5

See example P\_06 in the *Graphical Mode Examples Guide*.

#### **TTT diagram**

Use a TTT-diagram calculation type to do a precipitation simulation of the time-temperaturetransformation (TTT).

Temperature: Enter Min, Max, and Step values for the and choose a Unit (Kelvin, Celsius, or Fahrenheit).

Max annealing time: Enter a numerical value and choose a Unit (Seconds, Minutes, Hours, or Days).

Stop criteria: Choose Volume fraction of phase or % of equilibrium fraction and then enter a numerical value in the field. For each temperature, the simulation stops when the stop criteria is fulfilled or if the maximum simulation time is reached, whichever happens first.



See example P\_03 in the *Graphical Mode Examples Guide*.

### **CCT diagram**

Use a **CCT-diagram** calculation type to do a precipitation simulation of the continuous-coolingtransformation (CCT)

Temperature: Enter Min and Max values and choose a Unit (Kelvin, Celsius, or Fahrenheit).

Cooling rate(s): Enter a range of values in the field, e.g. .01 .1 1 10 100. These values are equal to K/s, °C/s or °F/s per second based on the Unit selected above.

Stop criteria: Enter a numerical value for the Volume fraction of phase. For each temperature, the simulation stops when the stop criteria is fulfilled or if the maximum simulation time is reached, whichever happens first.



See example P 07 in the *Graphical Mode Examples Guide*.

#### Numerical Parameters



"Numerical Method" on page 83 or search the online help.



Click **Options** to go to these settings.

#### Max time step fraction

The maximum time step allowed for time integration as fraction of the simulation time. The default is 0.1.

#### No. of grid points over one order of magnitude in radius

Default number of grid points for every order of magnitude in size space. The default is 150.0.

#### Max no. of grid points over one order of magnitude in radius

The maximum allowed number of grid points in size space. The default is 200.0.

#### Min no. of grid points over one order of magnitude in radius

The minimum allowed number of grid points in size space. The default is 100.0.

#### Max relative volume fraction of subcritical particles allowed to dissolve in one time step

The portion of the volume fraction that can be ignored when determining the time step. The default is 0.01.

#### Max relative radius change

The maximum value allowed for relative radius change in one time step. The default is 0.01.

#### Relative radius change for avoiding class collision

Set a limit on the time step. The default is 0.5.

#### Max overall volume change

This defines the maximum absolute (not ratio) change of the volume fraction allowed during one time step. The default is 0.001.

#### Max relative change of nucleation rate in logarithmic scale

This parameter ensures accuracy for the evolution of effective nucleation rate. The default is 0.5.

#### Max relative change of critical radius

Used to place a constraint on how fast the critical radium can vary, and thus put a limit on time step. The default is 0.1.

### Min radius for a nucleus to be considered as a particle

The cut-off lower limit of precipitate radius. The default is 5.0E-10 m.

### Max time step during heating stages

The upper limit of the time step that has been enforced in the heating stages. The default is 1.0 s.

### Maximum relative solute composition change at each time step

Set a limit on the time step by controlling solute depletion or saturation, especially at isothermal stage. The default is 0.01.

# **Plot Renderer**

The following is information about the settings available for a Plot Renderer when it is a successor to a Precipitation Calculator.

"Plot Renderer" in the *Thermo-Calc User Guide* or search for it in the Online Help.

There are also specific settings related to non-isothermal simulations that are detailed in this topic.

# **Plot Settings**

#### **Legend option**

Select whether the diagram's legend displays **On** or **Off**.

#### **Axis variable**

Set the state variable you want plotted along the X-axis and the Y-axis.

Select the Separate multimodal PSD check box to enter settings for Points, Excess kurtosis, Valley depth ratio and Minimum peak.

See "Non-Isothermal Simulations" on page 54 for definitions.

#### Axis type

Select the type of axis scale: Linear, Logarithmic, Logarithmic 10 or Inverse.

#### Limits

Specify the range along the axis to show in the plot. In the fields, enter the minimum and maximum values of the axis variable. You can also determine the **step** size between the tick marks along each axis.

Select the **Automatic scaling** check box to allow the program to set the limits.

#### **Unit (***Time X Axis***)**

Choose a **Unit**: **Seconds**, **Hours**, or **Days**.

#### Add an axis and Remove this axis buttons

Use the Add an axis ② and Remove this axis ⑤ buttons to add additional X- and/or Y-axes to a plot or to merge plots (to overlay one plot with another in the same diagram). When merging plots, you define this variable as the Z-axis.

#### **Axis Variables**

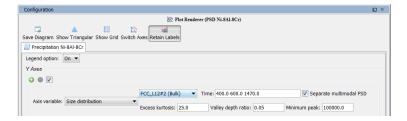
- **Mean radius** Spherical radius of average volume of all particles for a specific phase and nucleation type, regardless of their actual shapes.
- Critical radius Spherical radius of critical nuclei for a specific phase and nucleation type.
- Matrix composition Instantaneous compositions of the matrix phase.
- **Number density** Instantaneous number of particles per unit volume for a specific phase and nucleation type.
- **Size distribution** Number of particles varying with their sizes per unit volume per unit length, for a specific phase and nucleation type, at a specific time.
- **Volume fraction** Instantaneous volume fraction for a specific phase and nucleation type.
- **Nucleation rate** Instantaneous number of nuclei per unit volume per unit time for a specific phase and nucleation type.
- Driving force Instantaneous nucleation driving force per unit mole of elements for a specific phase and nucleation type. It is normalized with RT and therefore, dimensionless.
- **Mean cubic factor** Average cubic factor of cuboid particles. Available only when cuboid is selected as the **Morphology**.
  - The following settings area available as indicated when **Needle**, **Plate** or **Cuboid** are selected as the **Morphology** in the *Precipitate Phase* section on the Precipitation Calculator.
- **Cubic factor distribution** Variation of cubic factor with particle size at a specific time. Available with a cuboid morphology.
- **Mean aspect ratio** Average aspect ratio of non-spherical particles. Available with a needle or plate morphology. Note that this is always larger than 1, where 1 = a sphere.
- **Mean particle length** Diameter of non-spherical particles along the longer axis. Available with a needle or plate morphology.
- Aspect ratio distribution Variation of aspect ratio with particle size at a specific time.
   Available with a needle or plate morphology.

## **Non-Isothermal Simulations**

When doing non-isothermal simulations it is common that particles grow in different generations. This results in multi-modal size distributions. To correctly estimate the properties of these different generations of particles you need to separate the peaks of multi-modal distributions.

#### Separate multimodal PSD

When the **Separate multimodal PSD** check box is selected on a Plot Renderer activity for the Precipitation Calculator, the size distribution is evaluated at the given time steps and checked for multi-modal peaks. These are separated and used to calculate the specified property.



It can be applied on the following plot properties:

- Mean radius
- Size distribution
- Number density
- Volume fraction

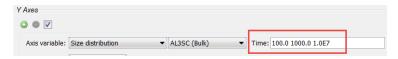
#### **Points**

This field is available when **Mean radius**, **Number density** or **Volume fraction** is selected as the **Axis variable**. Since the evaluation of multi-modality at each time step is costly, you can specify how many evaluation **Points** to use. The points are geometrically distributed over the time scale where nucleated particles are found. The default is 40 points.

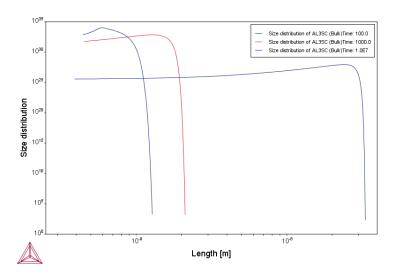
#### Time

This field is available when **Size distribution** is selected as the **Axis variable**. Enter one or a series of numbers in the field, separated by a space.

For example, if you enter 100.0, 1000.0 and 1.0E7 in the field:



When you click **Perform** the times are plotted:



#### **Excess Kurtosis**

The Excess kurtosis number specifies the sensitivity when the program evaluates whether a curve is bimodal or not. The evaluation process calculates the excess kurtosis of the given size distribution and compares it with the input value. If the value is lower than the given value, the distribution is split. The excess kurtosis is a statistical measure that measures how *peaked* and *how heavy tail* a distribution has. For the definition used see <a href="http://en.wikipedia.org/wiki/Kurtosis">http://en.wikipedia.org/wiki/Kurtosis</a>. The default is 25.0.

## Minimum Separation Limit (Valley Depth Ratio)

The **Minimum Separation Limit** is a rough method to prevent that noise or wiggles are interpreted as different distributions. If a local minima is encountered above this value the following peak is not interpreted as a new distribution. The valley depth ratio is the ratio of the minimum separation limit to the peak value. The default is 0.01.

### **Minimum Peak**

Any separated peak below the entered **Minimum peak** value is removed. The default is 100000.0.

# **Theoretical Models**

## In this section:

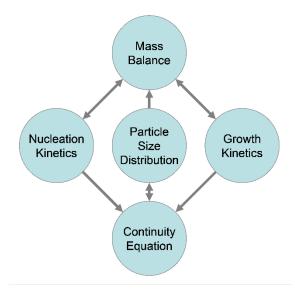
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# **Introduction to the Theory**

Based on the theory from Langer-Schwartz (1980)<sup>1</sup> Precipitation Module (TC-PRISMA) adopts Kampmann-Wagner numerical (KWN) (1991)<sup>2.</sup> method to simulate the concomitant nucleation. growth, and coarsening of precipitates in multicomponent and multiphase alloy systems. The KWN method is an extension of the original Langer-Schwartz (LS) approach and its modified (MLS) form, where the temporal evolution of the mean radius and particle density over the whole course of precipitation are predicted by solving a set of rate equations derived with certain assumptions for the rates of nucleation and growth, as well as the function of particle size distribution (PSD). The MLS approach differs from the LS with respect to the Gibbs-Thomson equations used for calculating equilibrium solubilities of small particles. The former applies the exact exponential form, whereas the latter takes the convenient linearized version. Instead of assuming a PSD function a priori and working with rate equations for determining only mean radius and particle density, the KWN method extends the LS and MLS approaches by discretizing the PSD and solving the continuity equation of the PSD directly. Therefore, the time evolution of the PSD and its  $n^{th}$  moment (0: number density;  $1^{st}$ : mean radius;  $3^{rd}$ : volume fraction) can be obtained altogether during the simulation. The key elements of the KWN method are the models for nucleation and growth under the mean field mass balance condition and the numerical algorithm for solving the continuity equation of the PSD. Coarsening comes out naturally without any ad hoc treatment.

The key elements of the KWN method and their relationship as implemented in the Precipitation Module (TC-PRISMA).



<sup>1.</sup> Langer, J. S., and Schwartz, A. J. (1980). "Kinetics of nucleation in near-critical fluids," *Physical Review A*, 21(3), 948–958.

<sup>&</sup>lt;sup>2.</sup> 'Homogeneous second phase precipitation', in: R.W. Cahn, P. Haasen, E.J. Kramer (Eds.), *Mater. Sci. Technol.*, Wiley-VCH Verlag GmbH & Co. KGaA, Weinheim, Germany: pp. 213–304.

# **Integration of the Precipitation Module into Thermo-Calc**

Precipitation Module (TC-PRISMA) is integrated with Thermo-Calc in order to directly get all necessary thermodynamic and kinetic information required in the KWN method. For industry relevant multicomponent alloys, thermodynamic and kinetic databases and calculation tools have to be used in order to obtain various quantities in the multicomponent models for nucleation and growth, such as the driving forces for the formation of embryos and their compositions, the atomic mobilities or diffusivities in the matrix, the operating interface compositions under local equilibrium conditions, the Gibbs-Thomson effect, and the deviation from local equilibrium due to interface friction etc. With Thermo-Calc and the Diffusion Module (DICTRA) as well as the accompanying databases, all these properties and effects can be calculated without unnecessary and inaccurate approximations.

In the following topics, various models and numerical methods implemented in Precipitation Module (TC-PRISMA) are introduced. Unless specified, spherical particles are assumed in the discussion.

# **Nucleation Theory**

Precipitation starts from the nucleation of clusters that can be considered as embryos of new phases with distinctive structures or compositions. In a perfect single crystal, nucleation happens homogeneously. In an imperfect crystal or polycrystalline materials, nucleation tends to occur heterogeneously due to the presence of dislocations, grain boundaries, grain edges, and grain corners. These imperfections or defects reduce the nucleation barrier and facilitate nucleation. However, if supersaturation or driving force is very large homogeneous nucleation is also possible since all sites including those inside a grain can be activated.

The following sections further elaborate on this theory.

- "Homogeneous Nucleation" on the next page
- "Heterogeneous Nucleation" on page 68
- "Nucleation During a Non-isothermal Process" on page 74

# **Homogeneous Nucleation**

The classic nucleation theory (CNT) [Kashchiev  $(2000)^1$  and Russell  $(1980)^2$ ] has been extended for the purpose of modeling nucleation in multicomponent alloy systems. The time dependent nucleation rate J(t) is given by

$$J(t)=J_sexpigg(rac{- au}{t}igg)$$

where

- ullet  $J_s$  is the steady state nucleation rate,
- $oldsymbol{ au}$  is the incubation time for establishing steady state nucleation conditions, and
- $oldsymbol{\cdot}$   $oldsymbol{t}$  is the time.

The steady state nucleation rate  $J_s$  is expressed by

$$J_s = Z eta^* N_0 expigg(rac{-\Delta G^*}{kt}igg)$$

where

- $oldsymbol{\cdot}$   $oldsymbol{Z}$  is the Zeldovich factor,
- $oldsymbol{eta^*}$  is the rate at which atoms or molecules are attached to the critical nucleus,
- ullet  $N_0$  is the number of available nucleation sites (of a particular type) per unit volume,
- ullet  $\Delta G^*$  is the Gibbs energy of formation of a critical nucleus,
- $oldsymbol{\cdot}$   $oldsymbol{k}$  is Boltzmann's constant, and
- ullet T is absolute temperature. In the case of homogeneous nucleation, each atom in the mother phase is a potential nucleation site.

The Gibbs energy of formation of a critical nucleus is expressed as

<sup>&</sup>lt;sup>1.</sup> Kashchiev, Dimo. 2000. Nucleation. Butterworth-Heinemann.

<sup>&</sup>lt;sup>2.</sup> Russell, Kenneth C. 1980. "Nucleation in Solids: The Induction and Steady State Effects." Advances in Colloid and Interface Science 13 (3–4): 205–318.

$$\Delta G^* = rac{16\pi\sigma^3}{3{\left(\Delta G_m^{lpha
ightarroweta}/V_m^{eta}
ight)^2}}$$
[Eq. 3]

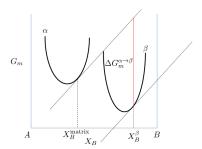
where

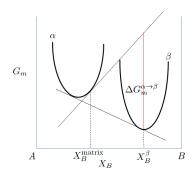
- $oldsymbol{\sigma}$  is the interfacial energy,
- $\Delta G_m^{\alpha o eta}$  is the molar Gibbs energy change for the formation of the eta precipitate of the critical composition  $X_i^{eta}$  from the  $\alpha$  matrix, i.e. the maximum driving force for the lpha o eta phase transformation
- $V_m^{eta}$  is the molar volume of the eta precipitate phase.

If elastic strain energy is excluded,  $\Delta G_m^{\alpha o \beta}$  is the chemical driving force for nucleation. There are two ways to calculate the nucleation driving force  $\Delta G_m^{\alpha o \beta}$  and the critical composition  $X_i^{\beta}$ , as shown in Figure 1 for a prototype binary system. The standard calculation, shown in

<u>Figure 1 (a)</u>, is to obtain the maximum driving force by finding the parallel tangent lines or surfaces passing through the alloy composition of the matrix, which is a routine calculation in Thermo-Calc. The approximate calculation, shown in <u>Figure 1 (b)</u>, is performed by using the Gibbs energy difference corresponding to the equilibrium composition in the precipitate phase. It can be used when the standard calculation fails, mostly within a miscibility gap.

Figure 1: Methods to calculate the nucleation driving force  $\Delta G_m^{\alpha op \beta}$  and the critical composition  $X_i^{\beta}$ . The first figure is a standard calculation and the second an approximate calculation.





Nucleation is a stochastic process and can be considered formally as a diffusion process or Brownian movement in the particle size space. Supercritical nuclei with radius slightly larger than the critical radius have a probability of passing back across the free energy barrier and dissolve in the matrix. The Zeldovich factor  $\boldsymbol{Z}$  is a measure of such probability and is related solely to the thermodynamics of the nucleation process in

$$Z=rac{V_m^eta}{2\pi N_A r^{*2}}\sqrt{rac{\sigma}{kt}}$$

where  $N_A$  is the Avogadro number and  $r^st$  , the critical radius, is given by

$$r^* = -rac{2\sigma V_m^eta}{\Delta G_m^{lpha 
ightarrow eta}}$$

The atomic or molecular attachment rate  $\beta^*$  reflects the kinetics of mass transport in the nucleation process and is given by Svoboda et. al (2004)<sup>1</sup>.

$$eta^* = rac{4\pi r^{*2}}{a^4} \Bigg[ \sum_{i=1}^k rac{ig(X_i^{eta/lpha} - X_i^{lpha/eta}ig)^2}{X_i^{lpha/eta}D_i} \Bigg]^{-1}$$

where

- a is the lattice parameter,
- $X_i^{eta/lpha}$  and  $X_i^{lpha/eta}$  are the mole fractions of element i at the interface in the precipitate and matrix, respectively.
- ullet  $D_i$  is the corresponding diffusion coefficient in the matrix.

<sup>&</sup>lt;sup>1.</sup> Svoboda, J., F.D. Fischer, P. Fratzl, and Ernst Kozeschnik. 2004. "Modelling of Kinetics in Multi-Component Multi-Phase Systems with Spherical Precipitates." Materials Science and Engineering: A 385 (1–2): 166–74.

The incubation time for an isothermal reaction is

$$au=rac{1}{ heta Z^2eta^*}$$

where heta differs from one derivation to another. The value 2 from Feder et al. (1966) $^{1}$  is adopted.

# **Elastic Strain Energy**

Elastic strain energy,  $E_{\rm el}$ , reduces the nucleation driving force,  $\Delta G_m^{\alpha \to \beta}$  in Equation 3, hence affecting nucleation rate and nuclei size. It also changes the shape of the particle by competing with interfacial energy.

### **Cuboid Particle**

"Precipitation Morphology" on page 89

The numerical simulations by Onaka et al.  $(2003)^2$  indicate that the elastic strain energy is reduced almost linearly with increasing cubic factor  $\eta$ . Since the numerical simulations are computationally expensive, we assume that the elastic strain energy follows the linear relationship with  $\eta$ , and the elastic strain energy for spherical ( $\eta=1$ ) and cubic ( $\eta=\sqrt{2}$ ) are calculated based on Khachaturyan's approximation (1983, 2008) $^3$ .

$$E_{
m el} = rac{1}{2}(c_{11} + 2c_{12})\epsilon_0^2 V[A_1 + A_2]$$

where

- $\epsilon_0$  is the lattice misfit strain.
- $c_{11}, c_{12}, c_{44}$  are elastic constants in cubic system.
- V is particle volume
- ullet  $A_1$  and  $A_2$  are expressed as

$$A_1 = 2rac{c_{11}-c_{12}}{c_{11}} - 12rac{c_{11}+2c_{12}}{c_{11}}rac{c_{11}-c_{12}-2c_{44}}{c_{11}+c_{12}+2c_{44}}I_1$$

<sup>&</sup>lt;sup>1.</sup> Feder, J., K.C. Russell, J. Lothe, and G.M. Pound. 1966. "Homogeneous Nucleation and Growth of Droplets in Vapours." Advances in Physics 15 (57). Taylor & Francis Group: 111–78.

<sup>&</sup>lt;sup>2.</sup> Onaka, Susumu, Noriko Kobayashi, Toshiyuki Fujii, and Masaharu Kato. 2003. "Energy Analysis with a Superspherical Shape Approximation on the Spherical to Cubical Shape Transitions of Coherent Precipitates in Cubic Materials." Materials Science and Engineering: A 347 (1–2): 42–49.

<sup>&</sup>lt;sup>3.</sup> Khachaturyan, Armen G. 2013. "Habit Plane and Orientation Relations in Precipitates: Comparison with Experimental Data." In Theory of Structural Transformations in Solids, 299–305. Mineola, New York: Dover Publications, Inc.

$$A_2 = -54rac{c_{11} + 2c_{12}}{c_{11}}rac{\left(c_{11} - c_{12} - 2c_{44}
ight)^2}{\left(c_{11} + c_{12} + 2c_{44}
ight)\left(c_{11} + 2c_{12} + 4c_{44}
ight)}I_2$$

with

**Sphere** 

$$I_1 = \frac{1}{15}$$
  $I_2 = \frac{1}{105}$ 

Cubic

$$I_1 = 0.006931$$
  $I_2 = 0.000959$ 

# **Ellipsoidal Particle (Plate and Needle)**

▶ "Precipitation Morphology" on page 89

Since they are ellipsoidal shape, it is convenient to use Eshelby's theory (1957)<sup>1.</sup> and (1959)<sup>2.</sup> with a reasonable computational cost. The Eshelby's tensor can be calculated by simply performing a surface integral over a unit sphere

$$D_{ijkl}=-rac{abc}{4\pi}\int_0^\pi\int_0^{2\pi}\Omega_{ij}n_kn_lrac{\sin heta}{eta^3}d\phi d heta$$

where

a, b, c are ellipsoid axes.

 $n_i (i=1,2,3)$  are unit directional vector normal to the spherical surface and

$$eta = \sqrt{(a^2\cos^2\phi + b^2\sin^2\phi)\sin^2 heta + c^2\cos^2 heta}$$

For matrix phase with cubic crystal symmetry, we have for  $\Omega_{ij}(i,j=1,2,3)$  See A. G. Khachaturyan (1983) $^6$ .

$$\Omega_{ii}(ec{n}) = rac{c_{44} + (c_{11} - c_{44})(n_j^2 + n_k^2) + \xi(c_{11} + c_{12})n_j^2n_k^2}{c_{44}D(ec{n})}$$

$$\Omega_{ij}(ec{n}) = -rac{(c_{12}+c_{44})(1+\xi n_k^2)}{c_{44}D(ec{n})}n_in_j$$

<sup>&</sup>lt;sup>1.</sup> Eshelby, J. D. 1957. "The Determination of the Elastic Field of an Ellipsoidal Inclusion, and Related Problems." Proceedings of the Royal Society A: Mathematical, Physical and Engineering Sciences 241 (1226): 376–96.

<sup>&</sup>lt;sup>2.</sup> Eshelby, J. D. 1959. "The Elastic Field Outside an Ellipsoidal Inclusion." Proceedings of the Royal Society A: Mathematical, Physical and Engineering Sciences 252 (1271): 561–69.

where

$$\xi = \frac{c_{11} - c_{12} - 2c_{44}}{c_{44}}$$

$$D(\vec{n}) = c_{11} + \xi(c_{11} + c_{12})(n_1^2n_2^2 + n_1^2n_3^2 + n_2^2n_3^2) + \xi^2(c_{11} + 2c_{12} + c_{44})n_1^2n_2^2n_3^2$$

The Eshelby S tensor can then be calculated as

$$S_{ijmn} = -rac{1}{2}C_{lkmn}\left(D_{iklj} + D_{jkli}
ight)$$

The total strain  $\epsilon_{ij}$  at each location inside the particle is related to its transformation strain (eigenstrain)  $\epsilon_{ij}^*$  as

$$\epsilon_{ij} = S_{ijkl} \epsilon_{kl}^*$$

Combined with elastic stress calculated as

$$\sigma_{ij} = C_{ijkl} \left( \epsilon_{kl} - \epsilon_{kl}^* 
ight)$$

The elastic strain energy can be then obtained

$$E^{el}=-rac{1}{2}\sigma_{ij}\epsilon_{ij}^{st}V$$

with V the particle volume.

# **Spherical Approximation for Nuclei**

In the Precipitation Module, the user-input or calculated interfacial energy is assumed to be the coherent interfacial energy that applies to the habit plane,  $\sigma^{\rm sph}_{\rm coh}$ , consistent with the approximation made by the embedded interfacial energy model. When calculating the critical nuclei, the interfacial energy in Equation 3 is assumed to be that of a spherical particle with

constant specific interfacial energy  $\sigma^{sph}_{coh}$ . This is consistent with the fact that the nuclei tend to be spherical due to dominant interfacial energy. Interfacial energy penalty assuming a nucleus with pre-defined, and most likely large, aspect ratio is thus over-estimated, and has been found to shut down nucleation event abnormally.

Under spherical approximation, the elastic strain energy is calculated using spherical expression in Equation 8 for spherical and cuboidal particles, or by setting a=b=c in Equation 9 for ellipsoidal particles.

# **Heterogeneous Nucleation**

All equations remain the same for the calculation of heterogeneous nucleation rate within the framework of CNT, but the nucleation energy and available nucleation site are different.

To a first approximation, the nucleation energy is calculated by assuming an effective interfacial energy for each heterogeneous nucleation site.

# Non-Spherical Particles and the Effect of Wetting Angle

"Precipitation Module (TC-PRISMA) References" on page 92

Non-spherical particles have been considered for grain boundary precipitation. Three morphologies are implemented for grain boundary, grain edge and grain corner precipitation, respectively, as shown in <u>Figure 1</u>. The parameter that defines the deviation from spherical shape is wetting angle (or dihedral angle),  $\theta$ , as described in <u>Figure 2</u> and taken from Clemm and Fisher (1955)<sup>1</sup>.

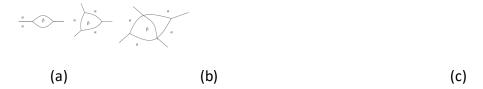


Figure 1: Non-spherical particles ( $\beta$ ) that precipitate at grain boundaries of matrix phase( $\alpha$ ) (a) grain boundary (two-grain junction) (b) grain edge (three-grain junction) (c) grain corners (four-grain junction). Images taken from Zang  $(2004)^2$ .

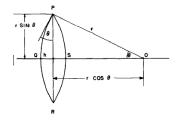


Figure 2: The nucleus at the grain boundary. Image from Clemm and Fisher, 1955.

Some physical parameters, mainly in nucleation models, have to be modified for the effect of the wetting angle. The following discussions contribute to this effect.

<sup>&</sup>lt;sup>1.</sup> Clemm, P J, and J C Fisher. 1955. "The Influence of Grain Boundaries on the Nucleation of Secondary Phases." Acta Metallurgica 3 (1): 70–73.

<sup>&</sup>lt;sup>2.</sup> Zang. 2004. 'Lecture 13: Heterogeneous Nucleation: Effects of Grain Boundaries and Surface Defects'. Lecture slides, Salt Lake City, Utah: The Zang Research Group, The University of Utah. Retrieved from http://www.eng.utah.edu/~lzang/images/lecture-13.pdf

# **Shape Factors**

Following Clemm and Fisher, the eliminated GB area between lpha-lpha grains due to the formation of nucleus of a new phase eta is:

[Eq. 1] 
$$A_{lphalpha}=ar^2$$

Where r is the radius of curvature of the nucleus. The surface area of the nucleus is:

[Eq. 2] 
$$A_{lphaeta}=br^2$$

and the volume of the nucleus is

[Eq. 3] 
$$V = cr^3$$

The expression of a, b, c in Equation 1, and 3 for grain boundary (two-grain junction), grain edge (three-grain junction) and grain corner (four-grain junction) can be found in the paper by Clemm and Fisher (1955)<sup>1</sup>.

# **Critical Radius and Activation Energy**

The energy change due to the formation of the new particle is thus

$$\Delta F = rac{\Delta G_m^{lpha 
ightarrow eta}}{V_m} \cdot c r^3 + b r^2 \sigma_{lpha eta} - a r^2 \sigma_{lpha lpha}$$

where  $\sigma_{lphaeta}$  and  $\sigma_{lphalpha}$  are the interfacial energy and grain boundary energy, respectively.

Then the critical radius should be

$$r^* = -rac{2ig(b\sigma_{lphaeta} - a\sigma_{lphalpha}ig)V_m}{3c\Delta G_m^{lpha o eta}}$$

And the activation energy barrier for nucleation is

$$W=rac{4}{27}rac{\left(b\sigma_{lphaeta}-a\sigma_{lphalpha}
ight)^3V_m^2}{c^2ig(\Delta G_m^{lpha
ightarroweta}ig)^2}$$

The interfacial energy, grain boundary energy and wetting angle can be related as

$$k=\cos heta=rac{\sigma_{lphalpha}}{2\sigma_{lphaeta}}$$
[Eq. 7]

i.e.

[Eq. 8] 
$$\sigma_{lphalpha}=2k\sigma_{lphaeta}$$

Which can be replaced into Equation  $\underline{5}$  and  $\underline{6}$ .

$$[ ext{Eq. 9}] ext{$r^*=-rac{2(b-2ak)\sigma_{lphaeta}V_m}{3c\Delta G_m^{lpha oetaeta}}$} \ W=rac{4}{27}rac{\sigma_{lphaeta}^3V_m^2}{\left(\Delta G_m^{lpha oeta}
ight)^2}rac{\left(b-2ak
ight)^3}{c^2}$$

The bulk, spherical precipitation equation is recovered by:

$$a=0,\quad b=4\pi,\quad c=rac{4\pi}{3}$$

so that

$$r^*=-rac{2\sigma_{lphaeta}V_m}{\Delta G_m^{lpha oeta}} \ W=rac{16\pi}{3}rac{\sigma_{lphaeta}^3V_m^2}{\left(\Delta G_m^{lpha oeta}
ight)^2}$$

And also the grain boundary precipitation with spherical shape (with weighting angle  $90^{\circ}$ ) follows Equations 12 and Equation 13 by  $k=\cos\theta=0$ .

### **Other Parameters**

#### **Zeldovich factor**

The Zeldovich factor is modified as

[Eq. 14] 
$$Z=Z_b\sqrt{f}$$

with  $Z_b$  the original value, and f the volume factor that is the ratio of the c in Equation 3 to the spherical shape factor  $4\pi/3$ .

$$f=rac{3c}{4\pi}$$

## **Impingement rate**

The surface area factor in impingement rate calculation is changed from  $4\pi(r^*)^2$  to  $b(r^*)^2$ .

## **Nucleation site density**

In each time step, the occupied grain boundary area

[Eq. 16] 
$$A_{\rm red} = a \cdot nr^{-2}$$

is deducted when calculating available nucleation site density.

#### **Growth rate**

The radius r defined in previous equations is the curvature of the precipitate surface (the Big radius) as shown in <u>Figure 2</u>. It is exactly the definition in the growth rate whose coarsening consideration relates to the curvature effect. Therefore, the growth rate equation can be directly used with r being defined here. Again, for bulk, spherical precipitation it automatically corresponds to the particle radius.

# **Output**

The volume of the precipitate should be calculated appropriately as  ${\it cr}^3$  with  ${\it r}$  being the curvature of the particle surface as defined above. For output of the particle size (mean radius, critical radius and particle size distribution) in the user interface, it seems better to use the radius of the eliminated grain boundary area, i.e.

$$r'=\sqrt{rac{a}{\pi}}\cdot r$$

With *a* defined in Equation 1.

For example, the size of the precipitate at grain boundary (two-grain junction) is the radius of the lens shown in Figure 2.

$$r'=r\sin heta=r\sqrt{1-k^2}$$

# The Shape and Size of Grains in the Matrix

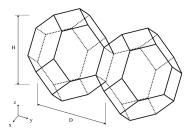


Figure 3: An example of tetrakaidecahedron approximation of grains

The number of different nucleation sites is dependent on the shape and size of grains in the matrix. Assuming all grains are equally large tetrakaidecahedra with the distance between one pair of square faces as H, and that between the other two pairs as D we obtain the densities  $\rho_2$ ,  $\rho_1$ , and  $\rho_0$  for the grain boundary area, edge length, and grain corner number respectively in

$$ho_2=rac{6\sqrt{1+2A^2}+1+2A}{4A}D^{-1}$$
  $ho_1=2rac{\sqrt{2}+2\sqrt{1+A^2}}{A}D^{-2}$   $ho_0=rac{12}{A}D^{-3}$ 

where A=H/D is the aspect ratio, defining the degree of elongation of the grains.

By default, the aspect ratio is set to 1 and the densities are then the same as obtained by Cahn (1956)<sup>1</sup>. Once the densities are known, if taking boundary thickness as one atomic layer, the available nucleation sites can be calculated by

$$N_i = 
ho_i igg(rac{N_A}{V_m^lpha}igg)^{i/3} \quad i=2,1,0$$

where  $V_m^lpha$  is the molar volume of the matrix phase and  $N_A$  is the Avogadro number.

For a crystalline material, given a dislocation density  $ho_d$ , the number of nucleation sites at the dislocations  $N_d$  can be calculated with the same form as in

<sup>&</sup>lt;sup>1.</sup> Cahn, John W. 1956. "Transformation Kinetics during Continuous Cooling." Acta Metallurgica 4 (6): 572–75.

$$N_d = 
ho_d igg(rac{N_A}{V_m^lpha}igg)^{1/3}$$

# **Nucleation During a Non-isothermal Process**

Under non-isothermal conditions, temperature dependency of key parameters such as nucleation driving force, solute diffusivities and solute concentrations, etc., have been taken into account, and are updated automatically during a simulation.

Another important parameter that depends on thermal history is the incubation time, defined by

$$au=rac{1}{ heta Z^2eta^*}$$

for an isothermal condition. In a non-isothermal process, the exact calculation of the incubation time requires a solution to the Fokker-Planck equation. In the Precipitation Module, an approximation approach has been employed to deal with the transient nucleation, which gives the incubation time as an integral form of past thermal history<sup>1.</sup> as in

$$\int\limits_{o}^{\tau}\beta^{*}(t')dt'=\frac{1}{\theta Z^{2}(\tau)}$$

where

au is the incubation time,  $oldsymbol{eta}^*$  is the impingement rate for solute atoms to the critical cluster as defined in

$$eta^*=rac{4\pi r^{*2}}{a^4}\Bigg[\sum_{i=1}^krac{ig(X_i^{eta/lpha}-X_i^{lpha/eta}ig)^2}{X_i^{lpha/eta}D_i}\Bigg]^{-1}$$

and Z is the Zeldovich factor, previously defined in

$$Z=rac{V_m^eta}{2\pi N_A r^{*2}}\sqrt{rac{\sigma}{kT}}$$

but now as a function of au derived from temperature change.

The starting point of the integral t'=0 is either the starting time if there is an initial nucleation driving force, or the latest time when the nucleation driving force is vanished.

<sup>&</sup>lt;sup>1.</sup> H.-J. Jou, P. Voorhees, and G.B. Olson (2004), 'Computer simulations for the prediction of microstructure/property variation in aeroturbine disks', *Superalloys*, 877–886.

#### Growth

## **Spherical Particles**

The growth rate models implemented in the Precipitation Module are called **Advanced** and **Simplified**.

The *Advanced* model is proposed by Chen, Jeppsson, and Ågren (CJA) (2008)<sup>1.</sup> and calculates the velocity of a moving phase interface in multicomponent systems by identifying the operating tie-line from the solution of flux-balance equations. This model can treat both high supersaturation and cross diffusion rigorously. Spontaneous transitions between different modes (LE and NPLE) of phase transformation can be captured without any *ad hoc* treatment. Since it is not always possible to solve the flux-balance equations and it takes time when possible, a less rigorous but simple and efficient model is preferred in many applications. The simplified model is based on the advanced model but avoids the difficulty to find the operating tie-line and uses simply the tie-line across the bulk composition.

All models treat a spherical particle of stoichiometric composition or with negligible atomic diffusivity growing under the local equilibrium condition.

According to the CJA model, the interface velocity v can be obtained together with interface concentrations by numerically solving 2n-1 equations, comprising of the flux balance equations for n-1 independent components and the local equilibrium conditions for all n components as in

$$v(c_i^{eta/lpha}-c_i^{lpha/eta})=c_i^{lpha/eta}M_irac{\mu_i^lpha-\mu_i^{lpha/eta}}{\xi_i r}$$
 [Eq. 2]  $\mu_i^{lpha/eta}=\mu_i^{eta/lpha}+rac{2\sigma V_m^eta}{r}$ 

where

- $c_i^{eta/lpha}$  and  $c_i^{lpha/eta}$  are the volume concentrations of component i at the interface in the precipitate and matrix, respectively,
- $M_i$  is the corresponding atomic mobility in the matrix,  $\mu_i^lpha$  and

<sup>&</sup>lt;sup>1.</sup> Chen, Qing, Johan Jeppsson, and John Ågren. 2008. "Analytical Treatment of Diffusion during Precipitate Growth in Multicomponent Systems." Acta Materialia 56 (8): 1890–96.

- $\mu_i^{lpha/eta}$  are the chemical potentials in the matrix of the mean-field concentration and at the interface, respectively.
- $\mu_i^{eta/lpha}$  is the chemical potential at the interface in the precipitate.

In the above local equilibrium condition, the multicomponent Gibbs-Thomson effect has been taken into account by adding a curvature induced pressure term to the Gibbs energy of the precipitate phase.

The introduced effective diffusion distance factor,  $oldsymbol{\xi_i}$  , for each independent component is given by

$$\xi_i = rac{\Omega_i}{2\lambda_i^2}$$

where

$$\Omega_i = rac{c_i^lpha - c_i^{lpha/eta}}{c_i^{eta/lpha} - c_i^{lpha/eta}}$$

is the so-called dimensionless supersaturation for an individual component, and  $\lambda_i$  is obtained via

$$2\lambda_i^2 - 2\lambda_i^3\sqrt{\pi}\exp{(\lambda_i^2)} ext{erfc}(\lambda_i) = \Omega_i$$

Combining Equation 1 and Equation 2, the simplified model is derived in

[Eq. 5] 
$$v = rac{K}{r} \Big[ \Delta G_m - rac{2\sigma V_m^eta}{r} \Big]$$

where

$$K=K_{
m sphere}=\left[\sum_irac{\left(X_i^{eta/lpha}(r)-X_i^{lpha/eta}(r)
ight)^2\xi_i}{X_i^{lpha/eta}(r)M_i}
ight]^{-1}$$

and  $\Delta G_m$  is the driving force.

The interface compositions  $X_i^{eta/lpha}(r)$  and  $X_i^{lpha/eta}(r)$  are simply taken from the tie line across the matrix composition.

The non-zero volume correction to the velocity according to Chen and Voorhees (1993)<sup>1.</sup> is taken into account and it follows as

$$v'=vig(1+r\sqrt{4\pi N_V\langle r
angle}ig)$$

where

- v' is the corrected velocity,
- ullet  $\langle m{r} 
  angle$  is the mean radius and
- $N_V$  is the number density.
- Also see "Coarsening" on page 80 for further discussion of these equations.

## **Non-Spherical Particles**

#### **Particle Shape Determination**

The shape of the cuboid particles, the cubic factor, is determined by the minimization of combined interfacial energy and elastic strain energy. If users choose transformation strain as **calculated from molar volume**, the lattice misfit is then calculated from difference in molar volume between matrix and precipitate phase, and the elastic strain energy is calculated accordingly. If users disregard the transformation strain, the elastic strain energy is neglected and hence the particles remain spherical shape.

If you choose **calculated aspect ratio from elastic energy** for plate and needle, the particle shape is determined by the minimization of combined interfacial energy and elastic strain energy. Otherwise, if you enter a value for aspect ratio, the particle shape is fixed except for nucleation calculation in which a spherical particle is always assumed,

"Homogeneous Nucleation" on page 62

#### **Interfacial Energy Anisotropy**

"Precipitation Morphology" on page 89

Isotropic interfacial energy is always assumed for spherical and cuboidal particles. For plate and needle, when the aspect ratio becomes larger than 1, the interfacial energy anisotropy occurs, so that the interfacial energy at the edge is larger than that at the broad face (habit plane). This increases the overall interfacial energy which is given as

<sup>&</sup>lt;sup>1.</sup> Chen, M K, and P W Voorhees. 1993. "The Dynamics of Transient Ostwald Ripening." Modelling and Simulation in Materials Science and Engineering 1 (5): 591–612.

**Plate** 

$$E_{
m int} = 4\pi \sqrt[3]{lpha^2} \sigma_0^{
m sph} r^2$$

Needle

$$E_{
m int} = 4\pi \sqrt[3]{lpha} \sigma_0^{
m sph} r^2$$

Where

- $oldsymbol{\alpha}$  is the aspect ratio
- $\sigma_0^{ ext{sph}}$  is the interfacial energy of the habit plane, i.e., the plane normal to the shorter axis
- $oldsymbol{\cdot}$  is the radius of a sphere with equivalent volume

#### **Growth Rate Adjustment**

"Precipitation Morphology" on page 89

For non-spherical particles, the growth rate equations for spherical particles are still applicable, but adjustment parameters are introduced to take into account the shape effect. The cuboid particles arise from "symmetry preserving" transformation, e.g., FCC\_A1 to  $\rm L1_2$ , and are thus highly isotropic and assumed growth rate equal to that of spherical particles. The plate and needle particles, on the other hand, arise from "symmetry breaking" transformations, e.g., cubic to tetragonal transformation, and are thus anisotropic leading to a significant increase of growth rate.

We define r as the radius of a sphere with equivalent volume of the non-spherical particle, so that the format of Equation 5 keeps unchanged. In our current model, two effects are considered contributing to the growth rate for plate and needle particles, from interfacial energy anisotropy and particle shape effect. The kinetic parameter K defined in Equation 5 is thus given as

[Eq. 8] 
$$K = K_{\sigma} \cdot K_{\mathrm{shp}} \cdot K_{\mathrm{sphere}}$$

with  $K_{\rm sphere}$  defined in Equation 6. The interfacial energy  $\sigma$  in Equation 5 is that of habit plane, i.e., the plane that is normal to the shorter axis of the particle.  $K_{\sigma}$  is the parameter that takes into account the Gibbs-Thomson effect due to interfacial energy anisotropy, based on Johnson (1965)<sup>1.</sup>

<sup>1.</sup> Johnson, Charles A. 1965. "Generalization of the Gibbs-Thomson Equation." Surface Science 3 (5): 429–44.

**Plate** 

$$K_{\sigma}=\sqrt[3]{lpha^2}$$

Needle

$$K_{\sigma}=\sqrt[3]{lpha}$$

where  $\alpha$  is the aspect ratio of the ellipsoidal particle.  $K_{\rm shp}$  is the parameter that takes into account the non-spherical concentration field around the particle. Assuming a shape-conserving concentration field and thus following the derivation by Ham(1958)<sup>1.</sup> and Ham(1959)<sup>2.</sup>, it is given as

#### **Plate**

$$K_{ ext{shp}} = rac{e\sqrt[3]{lpha}}{rccos(0) - rccos(e)}$$

Needle

$$K_{
m shp}=rac{2\sqrt[3]{lpha^2}e}{\ln(1+e)-\ln(1-e)}$$

where e is the eccentricity of the ellipsoidal particle.

$$e=\sqrt{1-rac{1}{lpha^2}}$$

<sup>1.</sup> Ham, Frank S. 1958. "Theory of Diffusion-Limited Precipitation." Journal of Physics and Chemistry of Solids 6 (4): 335–51.

<sup>&</sup>lt;sup>2.</sup> Ham, Frank S. 1959. "Shape-Preserving Solutions of the Time-Dependent Diffusion Equation." Quarterly of Applied Mathematics 17 (2): 137–45.

# Coarsening

Physically speaking, coarsening or Ostwald ripening where big particles grow and small particles shrink is a process driven by lowering the total surface energy of the system. From a thermodynamic point of view, the Gibbs-Thomson effect leads to inhomogeneous chemical potentials in the system if the particle sizes are not uniform. Solutes at the interface in the matrix near a particle of a radius smaller than critical radius have a higher chemical potential than that corresponding to the mean concentration of the matrix. As a result, the solutes diffuse from the precipitate/matrix interface to the inside of the matrix and cause dissolution of the particle. Conversely, particles with a radius larger than the critical size have lower interface potentials, and the solutes diffuse to the interface and cause growth of the particles.

Since it is not possible to find a closed form analytic solution for the problem of diffusion-controlled spherical particle dissolution (Aaron, 1970)<sup>1.</sup>, we simply apply the Growth equations 1 to 5 with the absolute value of  $\Omega_i$  to calculate the interface velocity for particles of all sizes.

As can be easily seen, if  $r < r^*$ , then the Gibbs-Thomson Equation 1 gives  $\mu_i^{\alpha/\beta} > \mu_i^{\alpha}$ , and a negative velocity results from Equation 2 for particles having  $r < r^*$ , which means that they shrink.

Results for particles having  $r>r^*$  are obtained vice versa. In all situations, when the absolute values of  $\Omega_i$  are very small, the steady-state solution for either growth or dissolution are recovered. In conclusion, the dissolution is treated as the reverse of growth (Aaron, 1970, Ibid.), and the coarsening comes out naturally either together with nucleation and growth or as a dominant process finally in the course of the evolution of the PSD.

<sup>&</sup>lt;sup>1.</sup> Aaron, Howard B. 1970. "Diffusion-Limited Phase Transformations: A Comparison and Critical Evaluation of the Mathematical Approximations." Journal of Applied Physics 41 (11). AIP Publishing: 4404.

# **Continuity Equation**

Let f(r) be the PSD of a precipitate phase, N the number of particles per unit volume,  $\langle r \rangle$  the mean radius and  $\phi$  - the particle volume fraction, is expressed as

$$_{ ext{ iny [Eq. 1]}} \qquad N = \int_0^\infty f(x) dr$$

$$_{[Eq.~2]}$$
  $\langle r 
angle = \int_0^\infty r f(r) dr$ 

The time evolution of  $m{f(r)}$  follows the continuity as in Langer and Schwartz (1980) $^1$ .

$$rac{\partial f}{\partial t} = -rac{\partial}{\partial r}[v(r)f(r)] + j(r)$$

Where v(r) is the growth rate of a particle of size r, and j(r) is the distributed nucleation rate, which is defined by

$$_{[Eq.~4]}$$
  $J=\int_{rst}^{\infty}j(r)dr$ 

where  $oldsymbol{J}$  is the nucleation rate.

<sup>&</sup>lt;sup>1.</sup> Langer, J. S., and A. J. Schwartz. 1980. "Kinetics of Nucleation in near-Critical Fluids." Physical Review A 21 (3): 948–58.

## **Mass Conservation**

The matrix concentration is updated at each time step according to the law of mass conservation. If the alloy concentration, i.e. initial mole fraction of component i in the matrix phase is  $X_i^0$ , the new concentration  $X_i$  can be obtained from the following mass balance shown in

$$X_{i}^{0} = \left(1 - \sum_{p} \int\limits_{0}^{\infty} \frac{4\pi r_{p}^{3} f(r_{p})}{3V_{m}^{p}} dr_{p}\right) X_{i} + \sum_{p} \int\limits_{0}^{\infty} \int\limits_{0}^{t_{i}} \frac{4\pi r_{p}^{2} f(r_{p}, t) v(r_{p}, t)}{V_{m}^{p}} X_{i}^{p}(r_{p}, t) dt dr_{p}$$

where

 $X_i^p(r_p,t)$  is the mole fraction of element i at the interface in the precipitate phase p of particle size  $r_p$  at time t.  $f(r_p,t)$ ,  $v(r_p,t)$  and  $X_m^p$  are the PSD function, growth rate, and molar volume of the precipitate phase p, respectively.

 $t_j$  is the time at time step j.

Apparently, the new matrix concentration, and thus the updated supersaturation, nucleation rate, and interface velocity are all functions of the PSD function. This inevitably leads to the complex non-linear behavior of and great difficulty in solving the continuity equation.

## **Numerical Method**

Since it is impossible to have a general close form solution, the complex non-linear continuity equation (see "Continuity Equation" on page 81) of the PSD function has to be solved numerically. Different numerical methods can be found in literature to solve this equation, such as finite difference and method of characteristics. In all approaches, the continuous PSD is discretized into a sufficiently large number of size classes, and then integrated discretely in the time space. The Precipitation Module utilizes the method of characteristics, where the number of particles in each size class is kept constant unless particles of the same size have been nucleated.

An adaptive geometric grid allocation method has been used for particle size discretization since from nucleation to coarsening the particle size can span several orders of magnitude. In this approach, the new size grids are allocated evenly in a logarithmic scale and the number of grid points over one order of magnitude is kept almost constant during the whole process by class merging and splitting. The time step is controlled by several adjustable numerical parameters based on mechanistic quantities.

Summarized below are all numerical parameters used in this approach to control either the size grid distribution or time steps.

#### **Maximum time step fraction**

 $P_{dt}^{max}$ 

Maximum time step allowed for time integration as fraction of the simulation time.

#### Number of grid points over one order of magnitude in r

 $P_{Nr}$ 

Default number of grid points for every order of magnitude in size space. The number determines a default ratio between two adjacent grid points. When there is a need to create new grid points, such as nucleating at a new radius not covered by the current range of PSD, this default ratio is used to add these new radius grid points. A larger value of this parameter enforces a finer grid to allow better numerical accuracy. However, this also comes with performance penalty, since finer grid in the size space often requires smaller time step to resolve the calculations.

## Maximum number of grid points over one order of magnitude in r

 $P_{Nr}^{max}$ 

The maximum allowed number of grid points in size space. This parameter determines a lower bound limitation for the ratio of every two next nearest grid points in order to maintain adequate computational efficiency. When a ratio of two next nearest grid points is less than this limit, the middle grid point is removed and the corresponding size class merged with the two neighbouring ones.

## Minimum number of grid points over one order of magnitude in r

# $P_{Nr}^{min}$

The minimum allowed number of grid points in size space. This parameter determines an upper bound limitation for the ratio of every two adjacent grid points in order to maintain proper numerical accuracy. When a ratio of two adjacent grid points exceeds this limit, a new grid point is then inserted between the two adjacent grids to keep the required resolution.

#### Maximum relative radius change

 $P_r$ 

The maximum value allowed for relative radius change in one time step. This parameter limits the time step according to the following relation, which is controlled by the particle growth:

$$\Delta t \leq P_r imes r/(|dr/dt|)$$
 for  $r > r_{dt}$ 

where  ${}^{r}dt$  is a cut-off subcritical size defined by the next parameter. The growth rates of supercritical particles (with  ${}^{r}>{}^{r}c$ ) are always bounded, and there is a size class and the corresponding growth rate that controls the time step. The subcritical particles (with  ${}^{r}<{}^{r}c$ ), however, has a mathematical singularity (negative infinity) in growth rate as  ${}^{r}{}^{r}$  approaches 0. This means that the time step can become extremely small if applying the above criterion to very small subcritical particles. In open literature, several researchers have tried mathematical transformation to avoid this singularity. Unfortunately, the transformation also complicates the formulation of the models. The Precipitation Module implementation uses a simple approach to deal with this issue by defining a cut-off size  ${}^{r}dt$ . All the particles with  ${}^{r}<{}^{r}dt$  may disappear within one time step.  ${}^{r}dt$  is determined by the next input parameter.

# Maximum relative volume fraction of subcritical particles allowed to dissolve in one time step

 $P_{rdt}$ 

This parameter represents the portion of the volume fraction that can be ignored when determining the time step. It is used to calculate the cut-off subcritical size,  $r_{dt}$ , for the above time step control that allows a maximum relative radius changes for all particles:

$$(\int_0^{rdt} fr^3 dr)/(\int_0^{\infty} fr^3 dr) = P_{rdt}$$

## Relative radius change for avoiding class collision

 $P_{cc}$ 

For the supercritical particles, the growth rate is non-linear – usually, it first increases with r and then decreases after a certain size. In the region(s) with growth rate decreasing with r, it is possible that the smaller size grid point can catch up with the larger size grid, if the time step is not controlled. To prevent this from happening, an additional parameter,  $P_{cc}$ , can be used to set a limit on time step according to the following relation:

$$\Delta t \leq P_{cc} imes (r_{i+1}-r_i)/(dr_i/dt-dr_{i+1}/dt)$$

for

 $r_{i+1} > r_{i+1}$ 

and

 $dr_{i+1}/dt < dr_i/dt$ 

#### Maximum overall volume change

 $P_{
u}$ 

This parameter defines the maximum absolute (not ratio) change of the volume fraction allowed during one time step. This parameter is also used in controlling allowable variation in volume fraction due to the newly created particles within one time step. That is

$$\Delta t \leq P_V/10/(4\pi r_{eff}^3 J/3)$$

where  $r_{eff}$  and J are effective radius and nucleation rate, respectively.

## Maximum relative change of nucleation rate in logarithmic scale

 $P_{logJ}$ 

This parameter ensures accuracy for the evolution of effective nucleation rate. It sets a limit on time step so that the relative change of nucleation rate does not exceed the specified value, based on the information of previous step. That is

$$\Delta t \leq P_{logJ} imes \Delta t_{previous}/(|log(J_1/J_2)|)$$

where nucleation rate  $J_1$  and  $J_2$  occurs at the beginning and end of  $\Delta t_{previous}$  .

#### Maximum relative change of critical radius

#### $P_{rc}$

During the nucleation under high supersaturation, the critical radius can vary dramatically. Hence, this parameter can be used to place a constraint on how fast the critical radium can vary, and thus put a limit on time step:

$$\Delta t \leq P_{rc} imes \Delta t_{previous}/(|r_{c1}-r_{c2}/r_{rc1}|)$$

#### Minimum radius for a nucleus to be considered as a particle

#### $P_{rmin}$

The cut-off lower limit of precipitate radius. Particles with radius smaller than the value specified for this parameter are discarded. In reality, the particle cannot be smaller than an atom; hence, there is no reason to keep track of particles of unphysical sizes.

## Maximum time step during heating stages

#### $P_{htmax}$

The upper limit of the time step that has been enforced in the heating stages. The current algorithm may over-estimate the subsequent time increment when temperature is increased. It is thus required to reduce this value when the calculation terminates unexpectedly during or after a heating stage.

#### **Numerical Control Parameters Default Values**

Default value for numerical parameters that controls the size grid distribution and time step.

Parameter	Default value
$P_{dt}^{max}$	0.1

Parameter	Default value
$P_{Nr}$	200
$P_{Nr}^{max}$	300
$P_{Nr}^{min}$	100
$P_r$	0.01
$P_{rdt}$	0.01
$P_{cc}$	0.5
$P_{ u}$	0.001
$P_{logJ}$	0.5
$P_{rc}$	0.1
$P_{rmin}$	5e-10m
$P_{htmax}$	1.0s

# **Estimation of Coherent Interfacial Energy**

Interfacial energy is an important parameter used in precipitation simulations to calculate the rates of nucleation, growth/dissolution, and coarsening. The value of interfacial energy can vary dramatically (usually between 0.01 to  $2.0 \text{ J/m}^2$ ).

The extended Becker's model functions to estimate coherent interfacial energy by using thermodynamic data from existing CALPHAD thermodynamic databases:

$$\sigma_c = rac{n_s z_s}{N_A z_l} \Delta E_s$$

where  $\sigma_c$  is the coherent interfacial energy,  $n_s$  is the number of atoms per unit area at the interface,  $z_s$  is the number of cross bonds per atom at the interface,  $z_l$  is the coordination number of an atom within the bulk crystal lattice, and  $\Delta E_s$  is the energy of solution in a multicomponent system involving the two phases being considered [Becker, 1938].

<sup>&</sup>lt;sup>1.</sup> Becker, 1938. 'Die Keimbildung bei der Ausscheidung in metallischen Mischkristallen', Ann. Phys. 424, 128–140.

# **Precipitation Morphology**

As the spherical particle is the default morphology, for precipitations within the grain, three non-spherical shapes are considered: *cuboid*, *plate* and *needle*.

- "Growth" on page 75
- "Homogeneous Nucleation" on page 62
- "Configuration Settings" on page 42

#### **Cuboid**

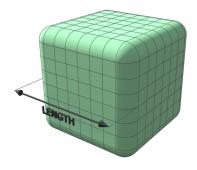


Figure 1: Cuboids have six faces, which form a convex polyhedron.

The cuboid shape is described by a supersphere

$$x_1^p+x_2^p+x_3^p=R^p \qquad (p\geq 2)$$

with p=2 being spherical shape. The larger the p, the more cubic the shape. Sometimes it is useful to define the cubical character as

$$\eta = \sqrt{2} \cdot 2^{-1/p}$$

Which gives a spherical shape when  $\eta=1$  , and a cubic shape when  $\eta=\sqrt{2}$ .

## **Plate**

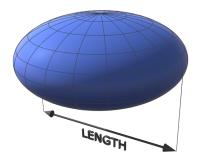


Figure 2: Oblate spheroids have rotational symmetry around an axis from pole to pole.

The plate is described as oblate spheroid

$$rac{x_1^2}{l^2} + rac{x_2^2}{l^2} + rac{x_3^2}{r^2} \leq 1 \qquad l > r$$

with aspect ratio

$$lpha=rac{l}{r}>1$$

#### Needle

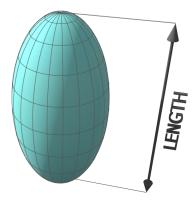


Figure 3: A prolate spheroid is a surface of revolution obtained by rotating an ellipse about its major axis.

The needle shape is described as prolate spheroid

$$rac{x_1^2}{r^2} + rac{x_2^2}{r^2} + rac{x_3^2}{l^2} \leq 1 \qquad l > r$$

with aspect ratio

$$lpha=rac{l}{r}>1$$

## Precipitations at a Grain Boundary, Edge or Corner

For precipitations at a grain boundary, grain edge or grain corner, non-spherical particles are considered based on wetting angle.

- > "Heterogeneous Nucleation" on page 68
- "Configuration Settings" on page 42

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